

BATSE – History

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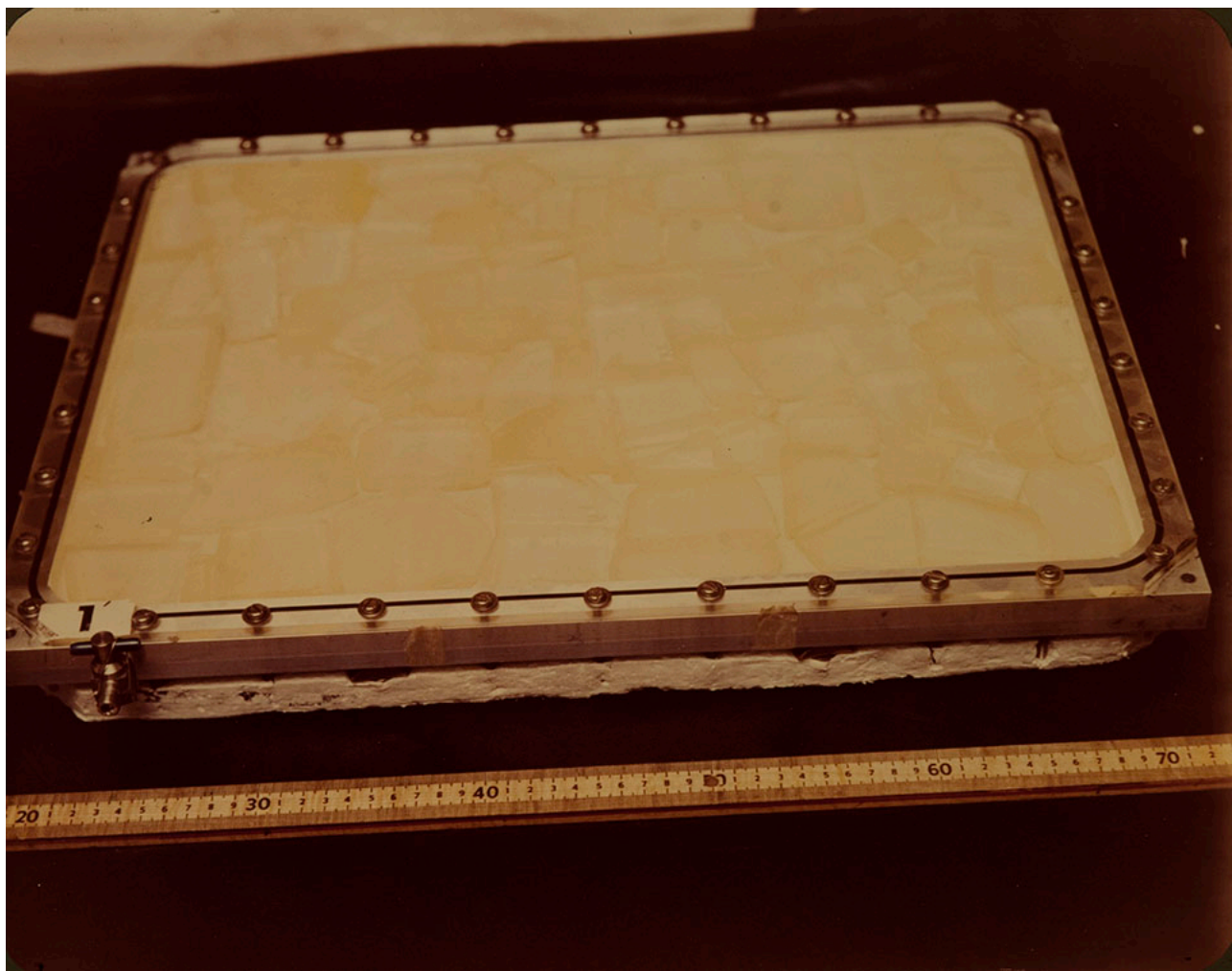
GRBs First Announced – 1973

- **Ap. J. Preprint of Discovery**
- **AAS Meeting – Columbus, Ohio**

**Huntsville “Group” (2) - Decision to build a large-area
detector balloon experiment**

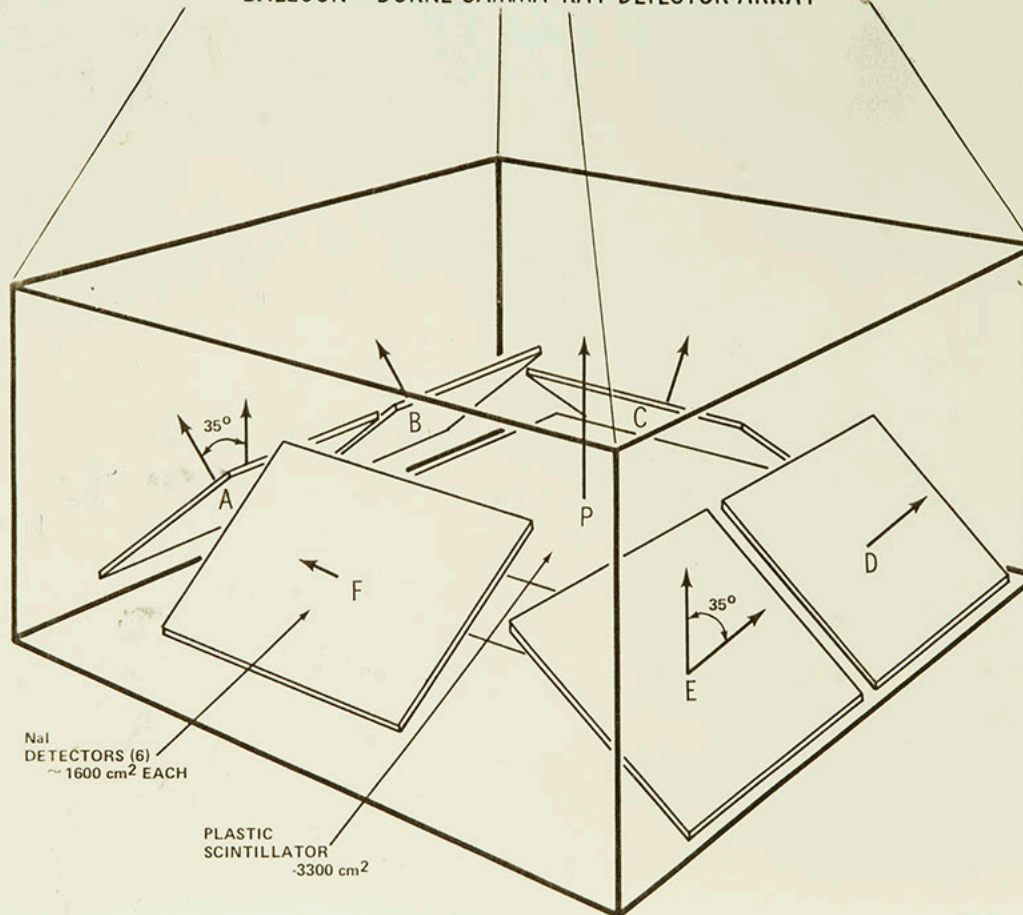
Detector Development & Early Balloon Flights

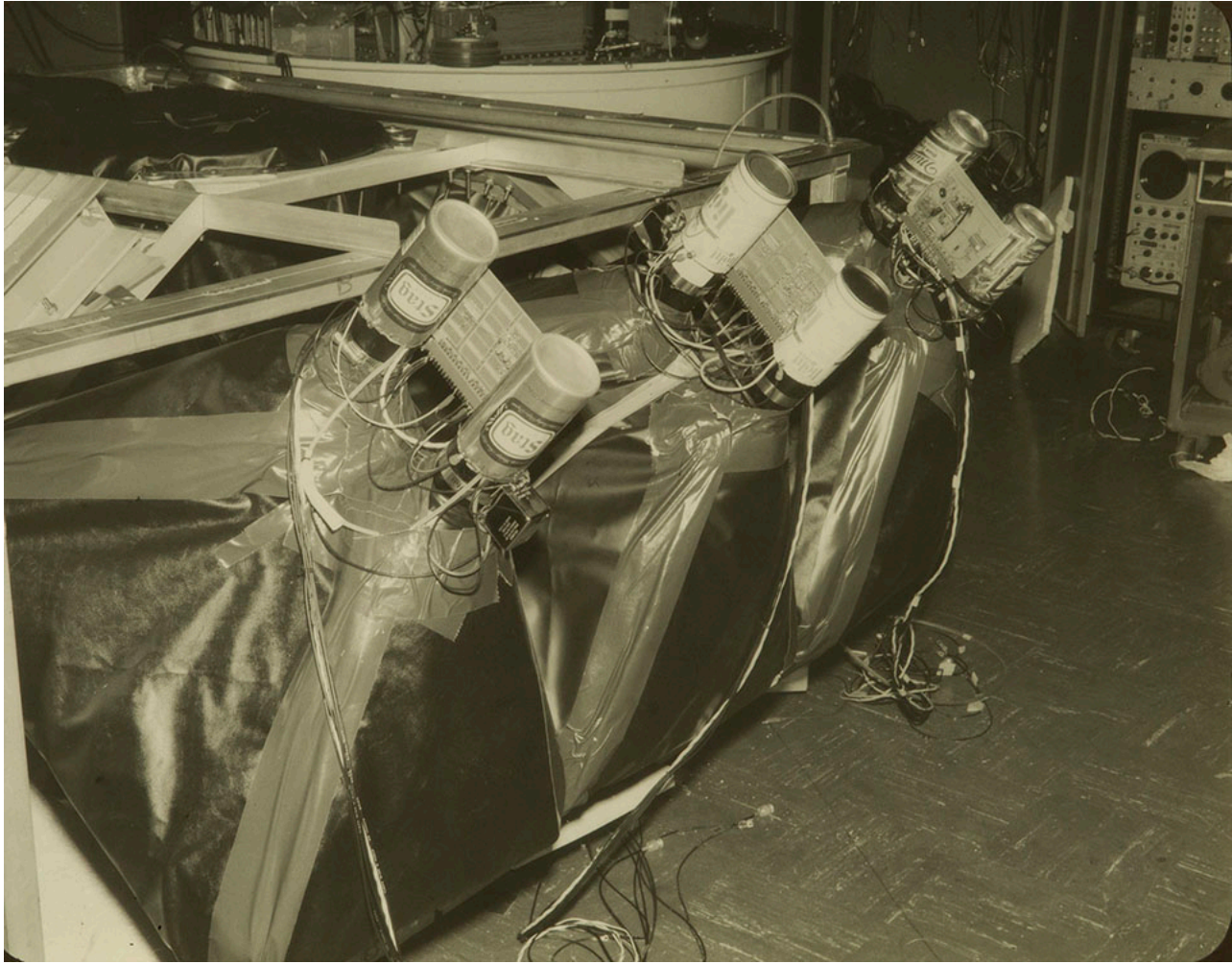
1974-1977





BALLOON - BORNE GAMMA-RAY DETECTOR ARRAY





1977-1978

GRO Announcement of Opportunity (AO)

- Proposed Transient Event Monitor (“TEM”)
- Twelve Detectors (Dodecahedron)
- Six on Top & Six on Bottom of S/C
- Partial Selection – Six Detectors
 - only on bottom of spacecraft

1979-1980

“TEM” - Name changed to BATSE

Negotiating after Approval:

- Requested 8 Detectors (for full-sky coverage)
(Octahedron - Four on top & bottom)**

GAMMA-RAY OBSERVATORY

CONCEPTUAL DESIGN

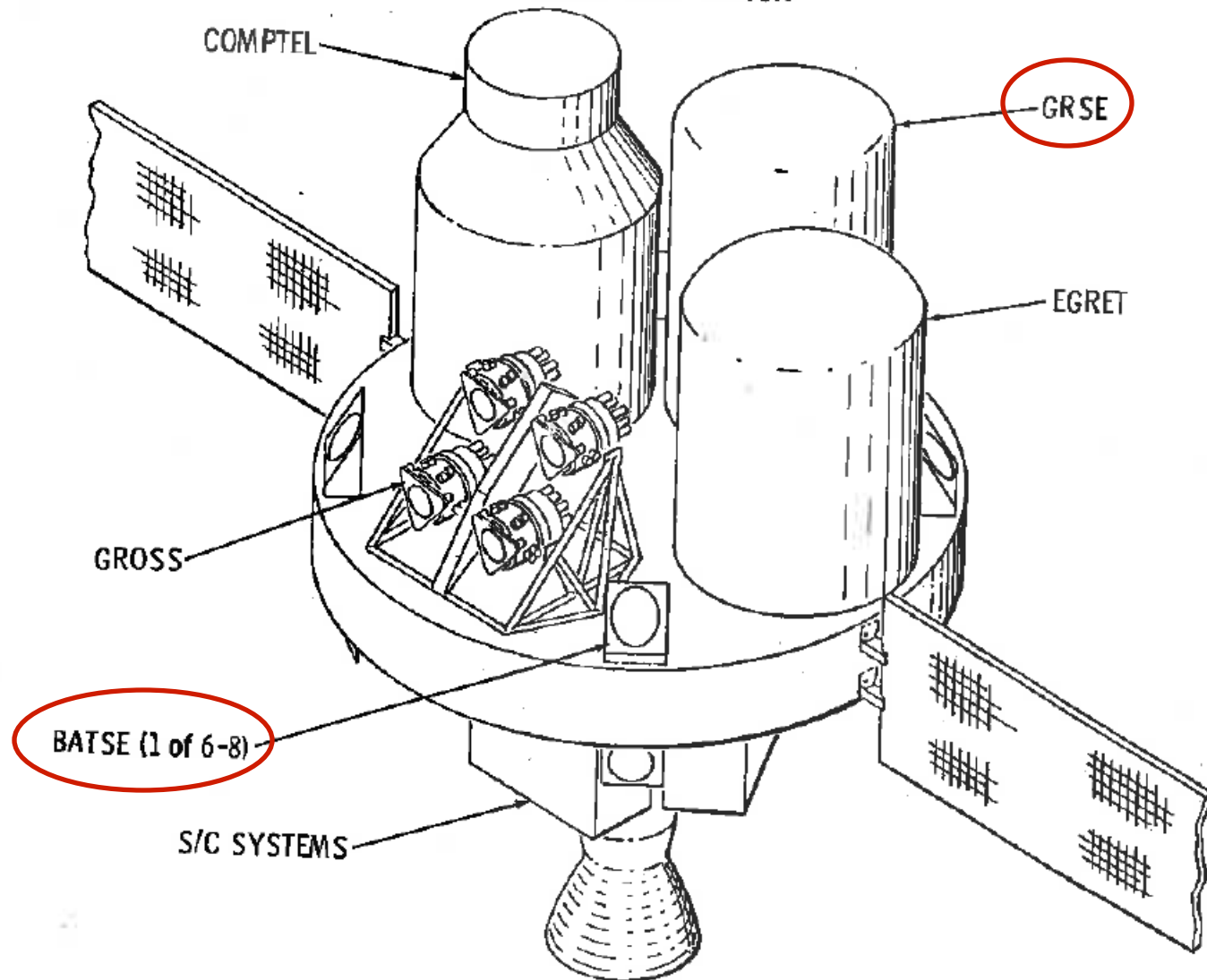
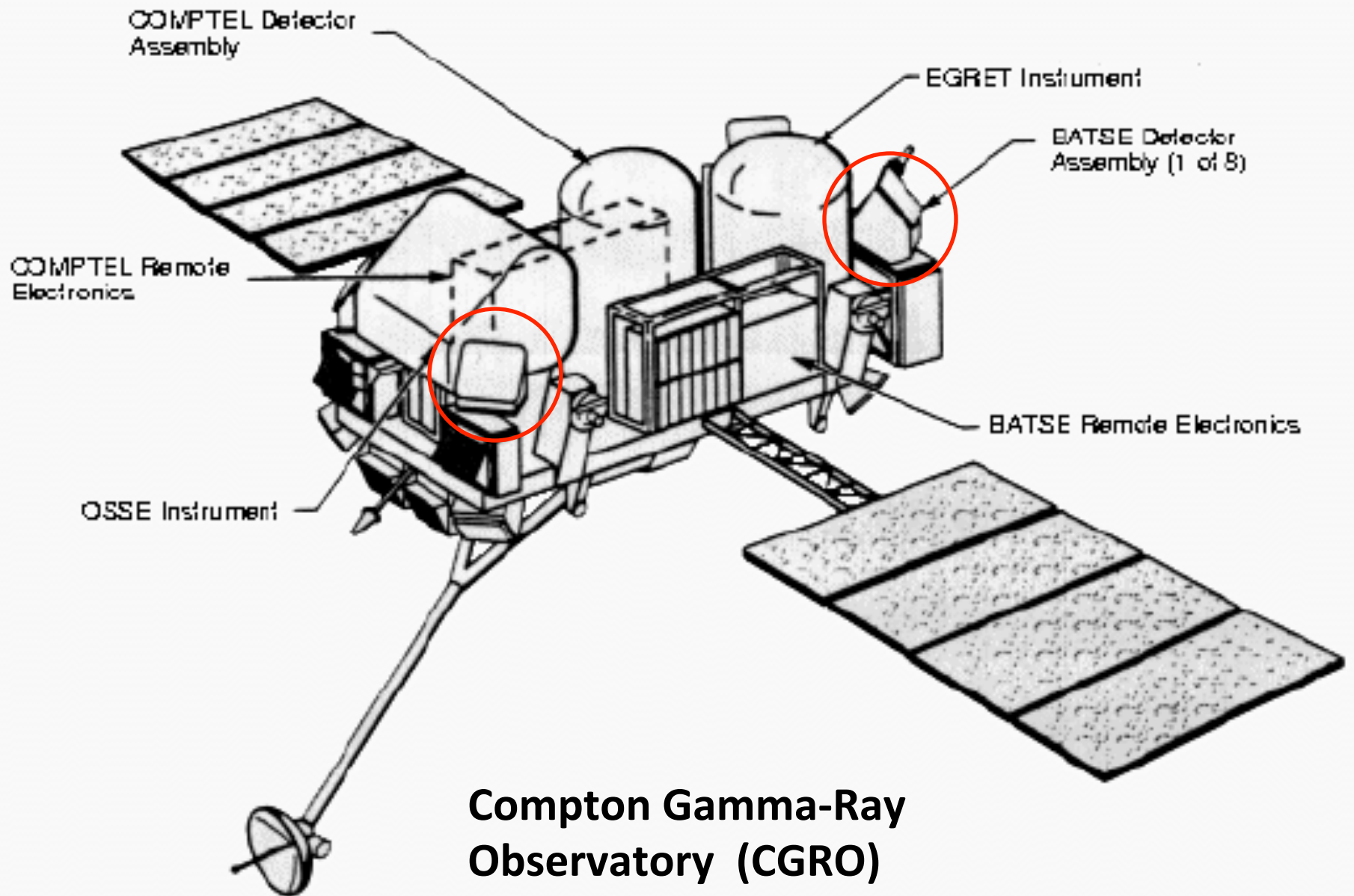


Fig. 3. Conceptual drawing of the planned Gamma-Ray Observatory (GRO) showing the four major pointed experiments and a possible configuration for the BATSE array.

1981-1982

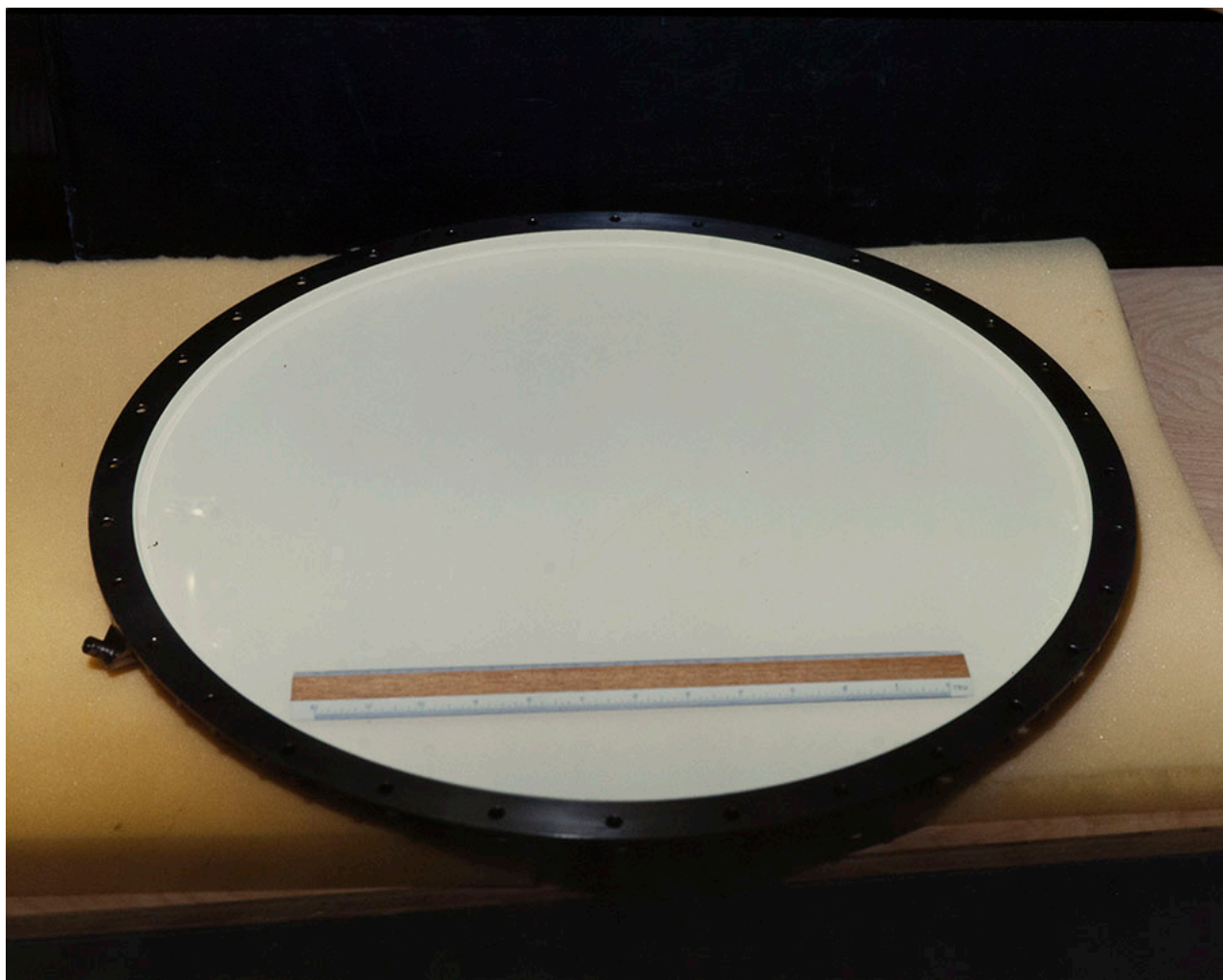
- GRSE Experiment Removed from GRO**
- Spectroscopy Detectors added to BATSE,
one in each detector module**
(a rare event; experiments are usually de-scoped by NASA)

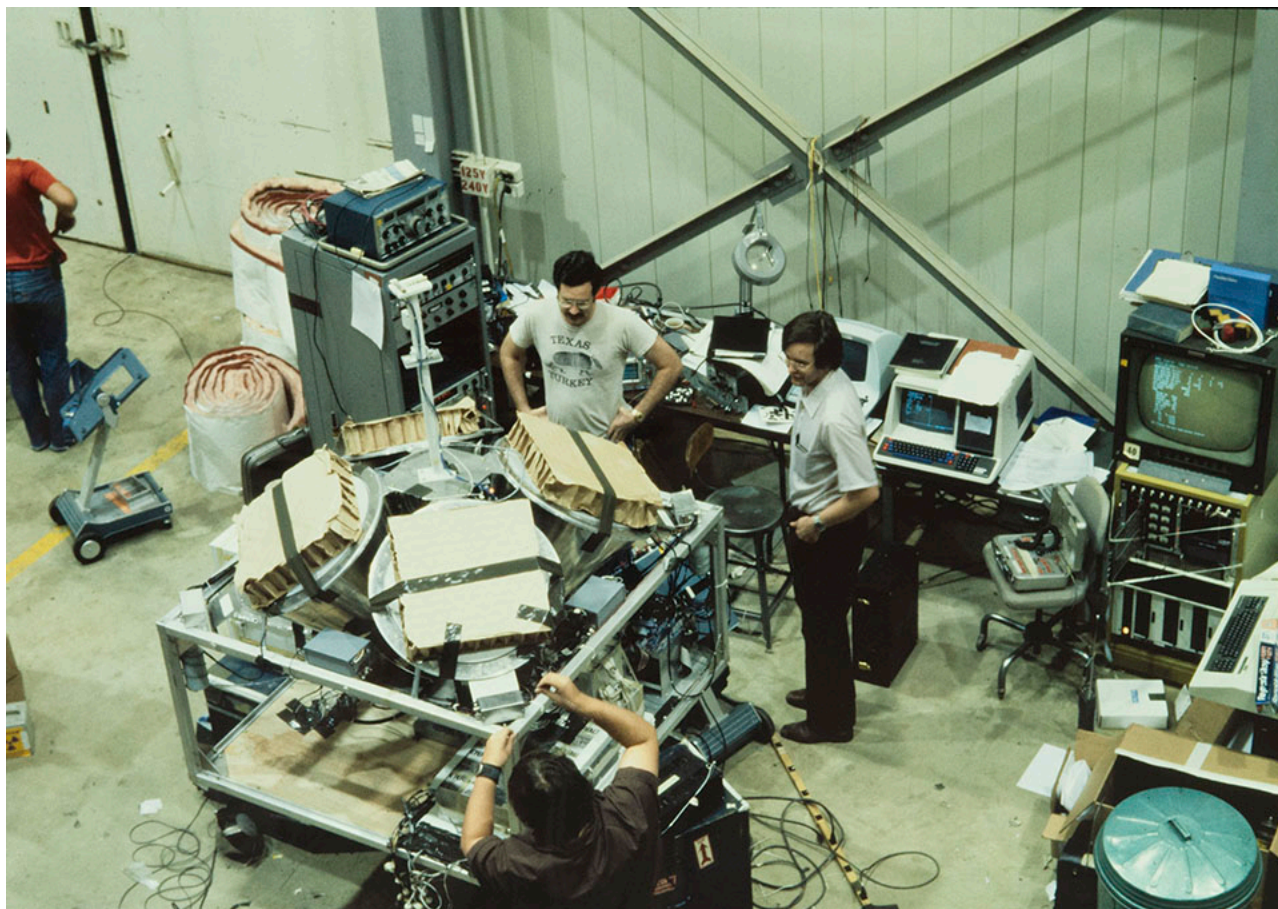


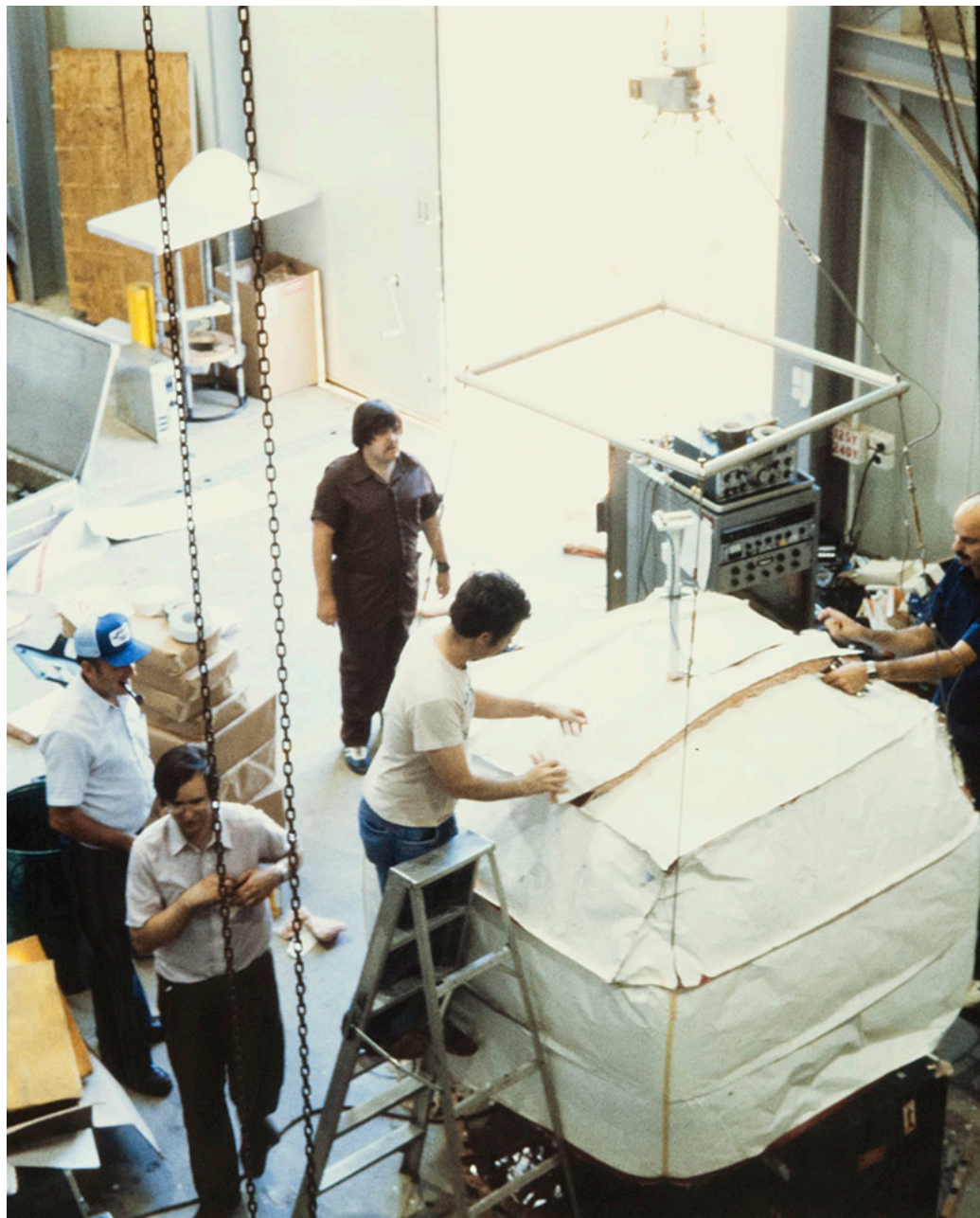
**Compton Gamma-Ray
Observatory (CGRO)**

BATSE Detector Development & Balloon Flights

1981 - 1987





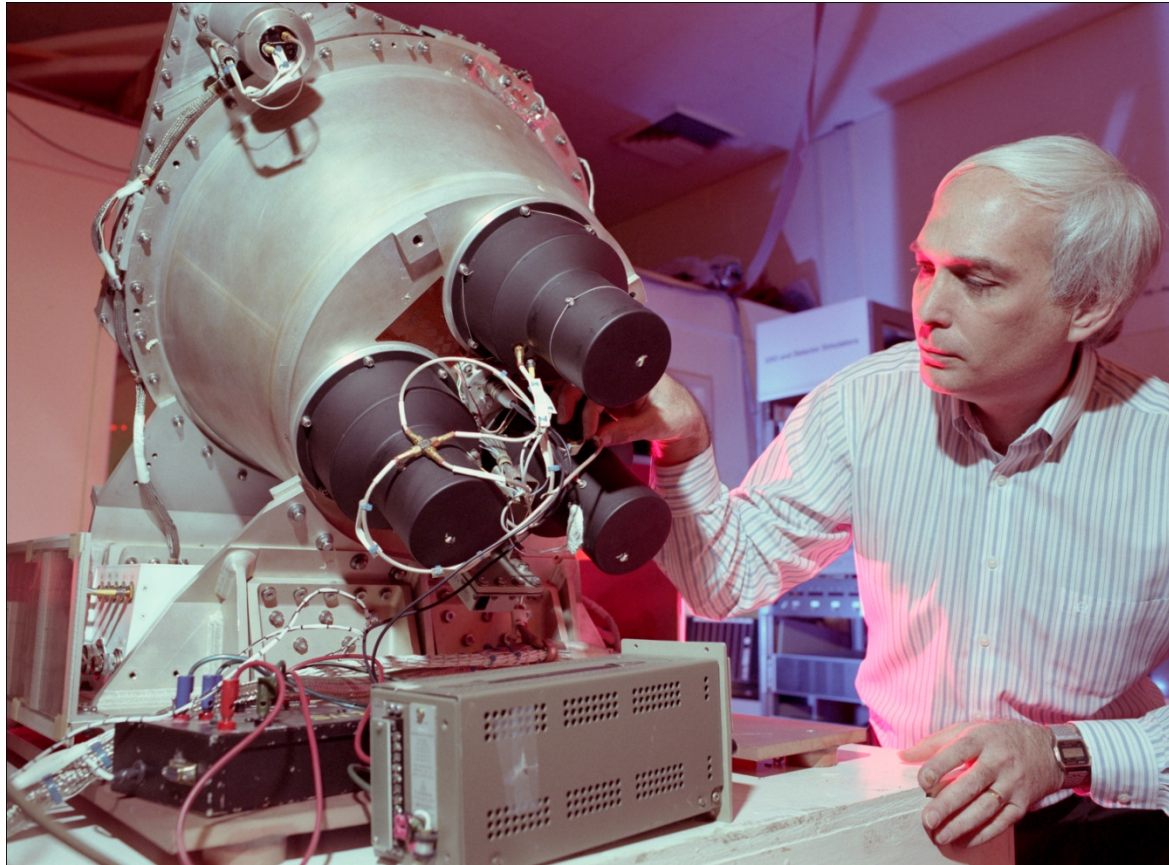




Building BATSE & GRO

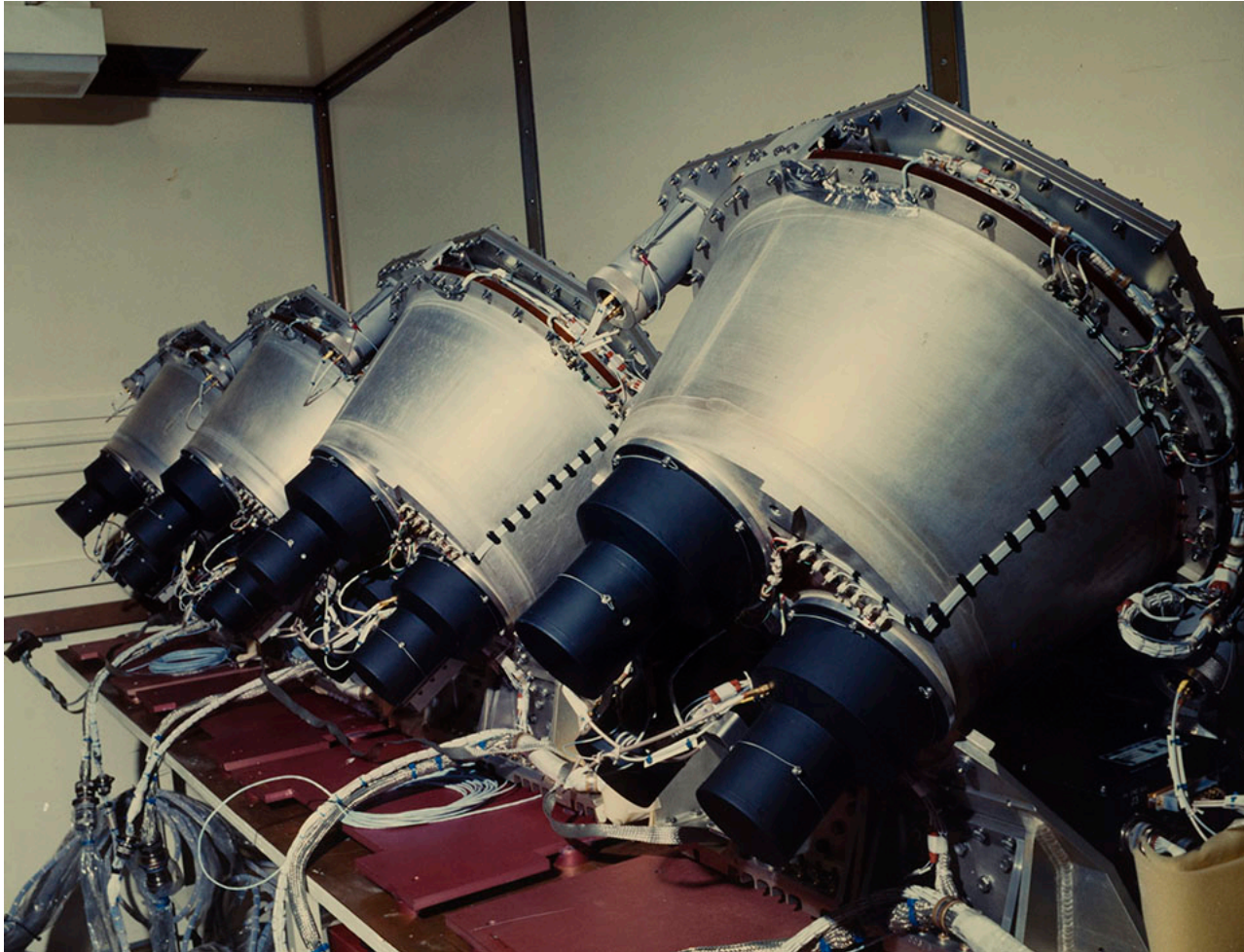
1982 - 1989

BATSE Detector Module



BATSE Detector Modules:

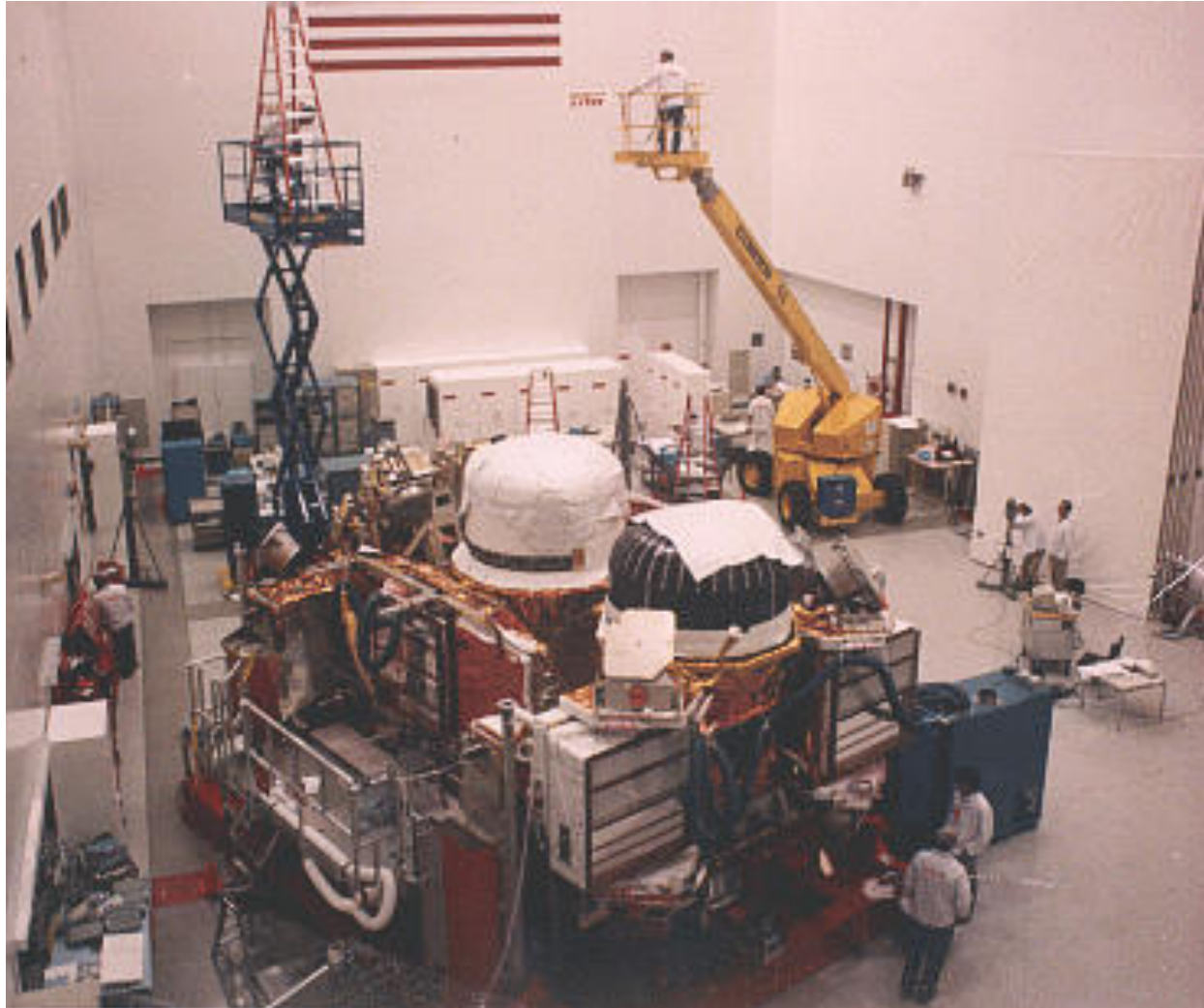
- Design, Fabrication, Testing & Calibration
at NASA-MSFC**



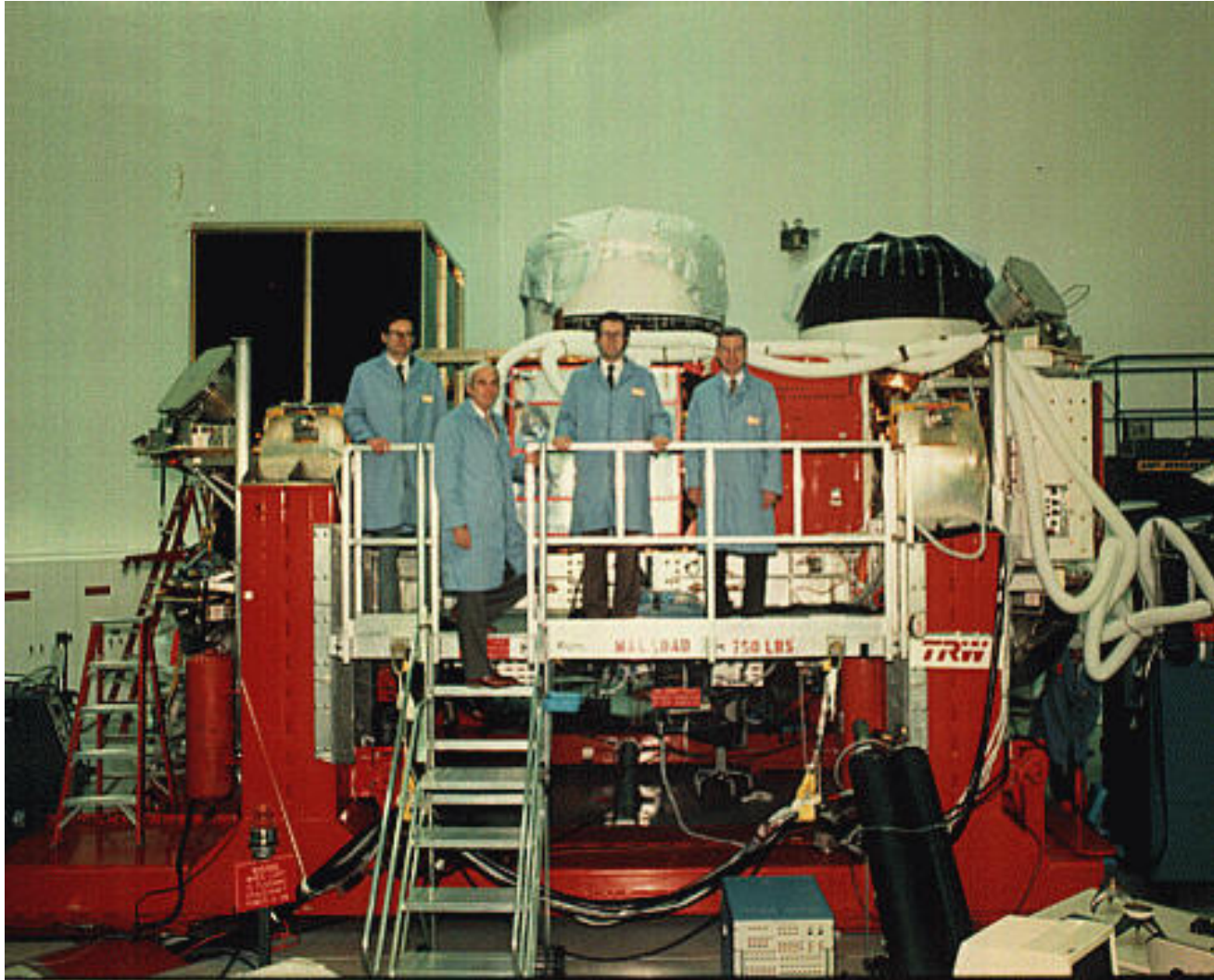
Integration of Experiments on GRO Spacecraft in California



Radioactive Source Survey of BATSE on the GRO Spacecraft



GRO P.I.s Near the GRO Instruments



Launching & Deploying GRO

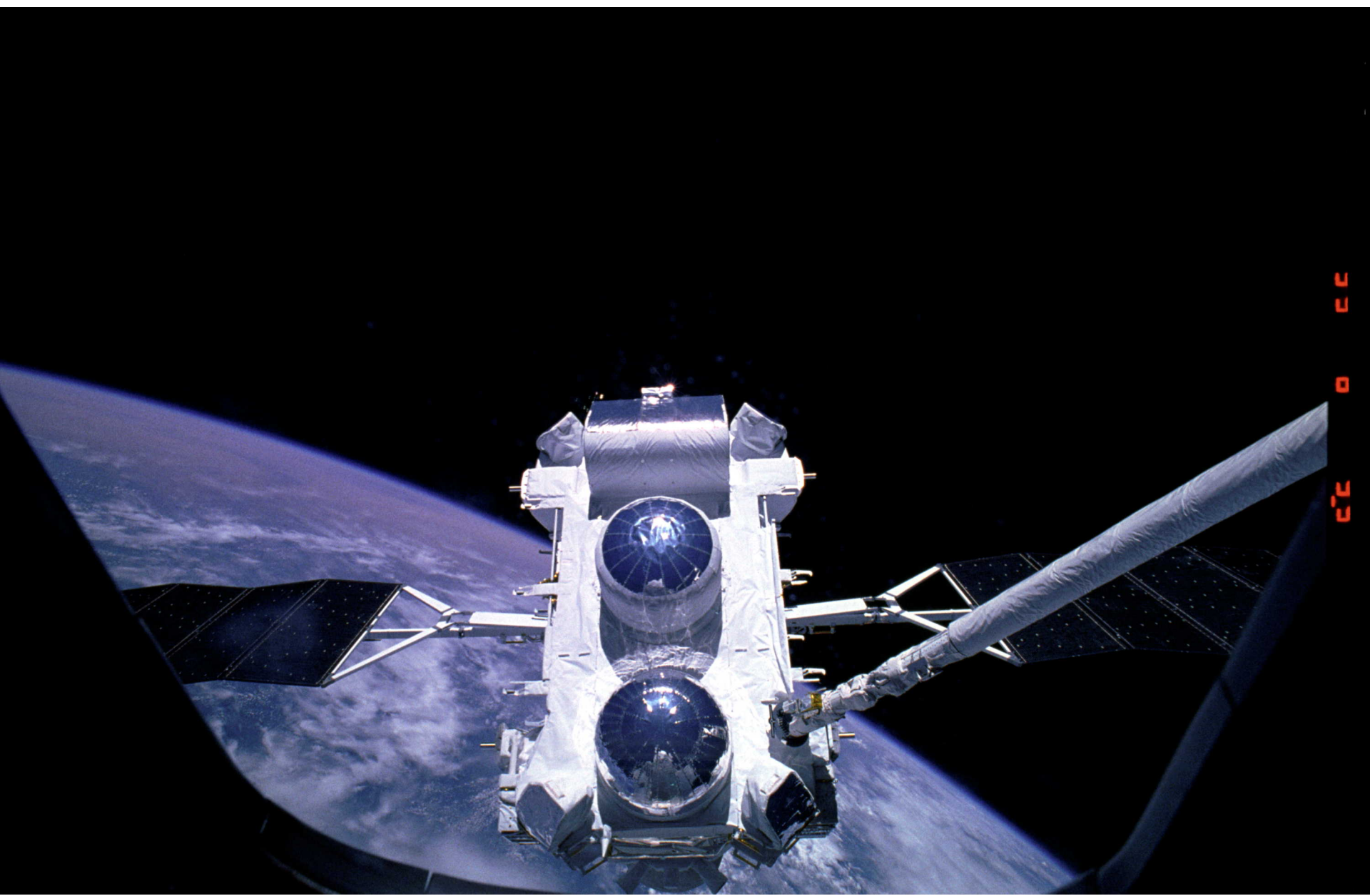
April 4-5, 1991

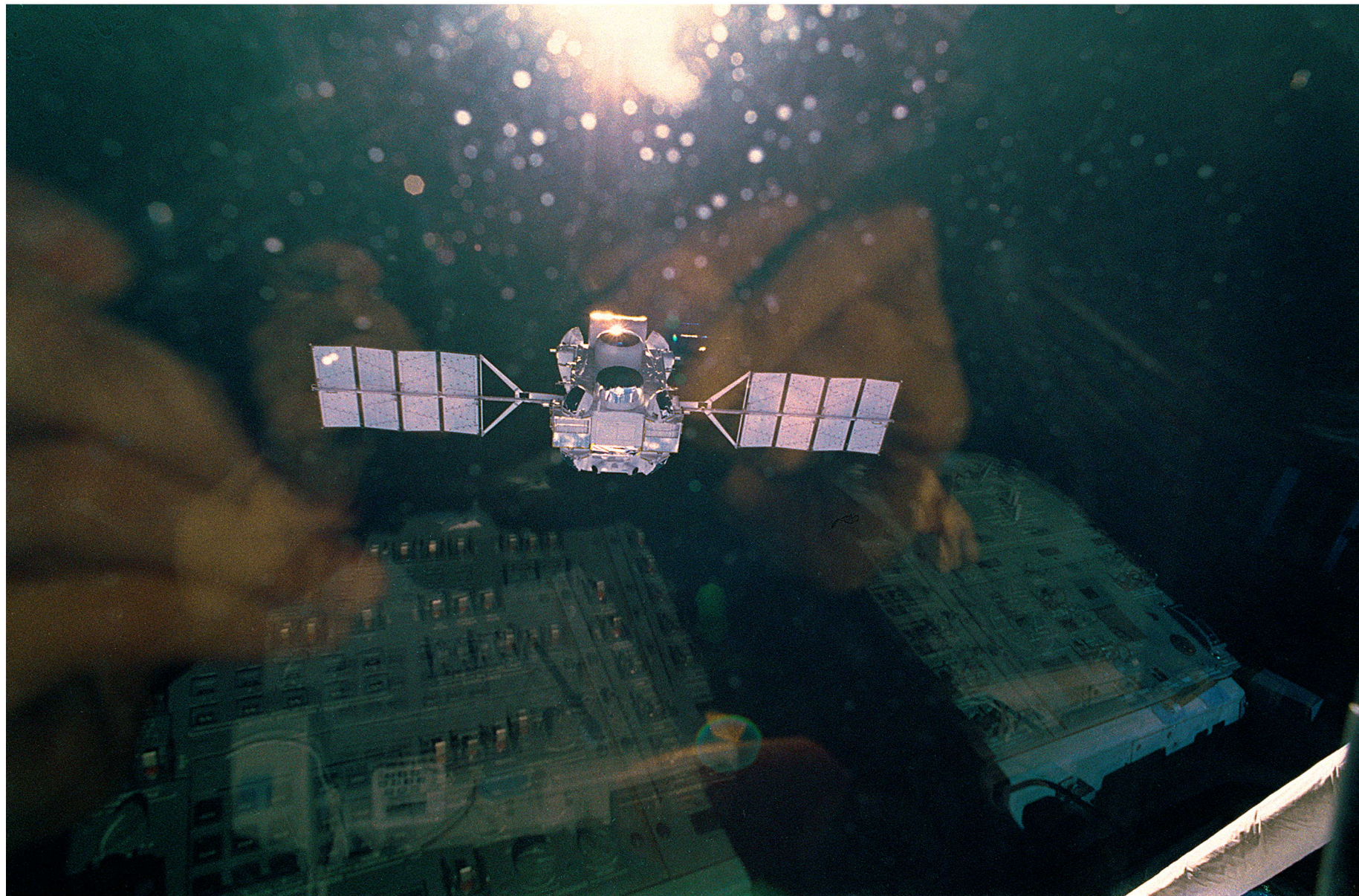
**Shuttle launch –
Gamma-Ray
Observatory (GRO)**

**(Operational:
re-named CGRO)**









nature

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MAPPING COSMIC GAMMA-RAY BURSTS

An inventory for the origins of life
More pulsar planets
Atomic structure of bacteriophage ϕ X174

BATSE

**First large experiment designed for
GRB studies:**

- **Full-sky**
- **Large area**
- **Good sensitivity for E_{peak} of most GRBs**
- **Moderate spectral capabilities** – (good enough for most time/spectral correlations and time resolved spectral studies)

BATSE Legacy

- Largest sample (2704 GRBs); full-sky, 9+ years in operation
- Well- characterized, full-sky instrument
- Likely will not be exceeded for several decades
- Led to GCN Network (S. Barthelmy)

>1000 Papers Based on BATSE Observations

>50 Ph.D. theses

BATSE - Major GRB Results

- **Global properties of GRB Distributions:**
 - Intensity Distribution & Sky Distribution
 - Not consistent with any Galactic Distribution, nearby extragal objects, incl. large clusters
 - Strong Indications that GRBs were at cosmological distances (Although BeppoSAX nailed it)
- **Comprehensive Temporal/Spectral Studies**
- **Two Populations of GRBs**
- **Rapid GRB Response:** beginning of Bacobine/GCN - led to breakthrough wide-field observations (e.g. GRB990123)

The End

Back-up Notes:

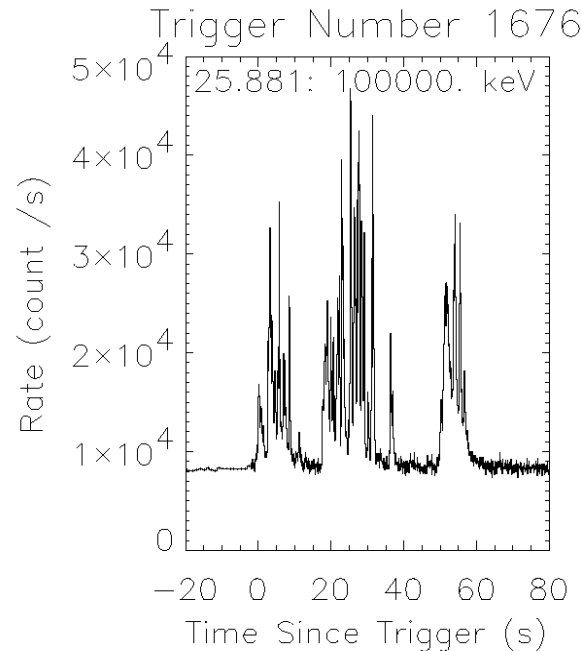
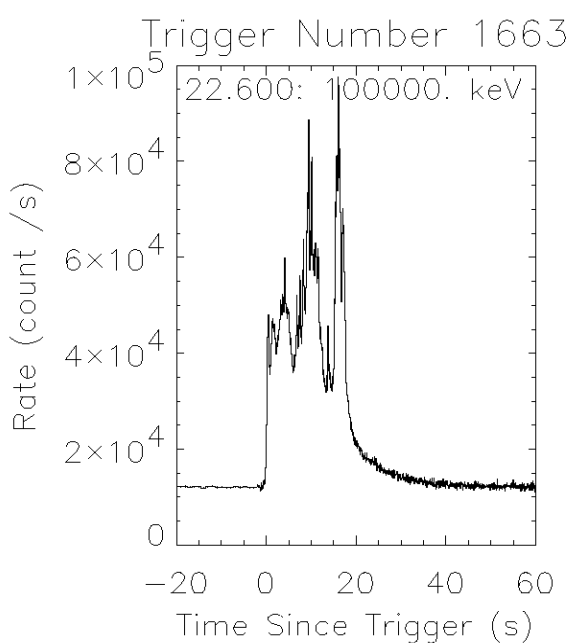
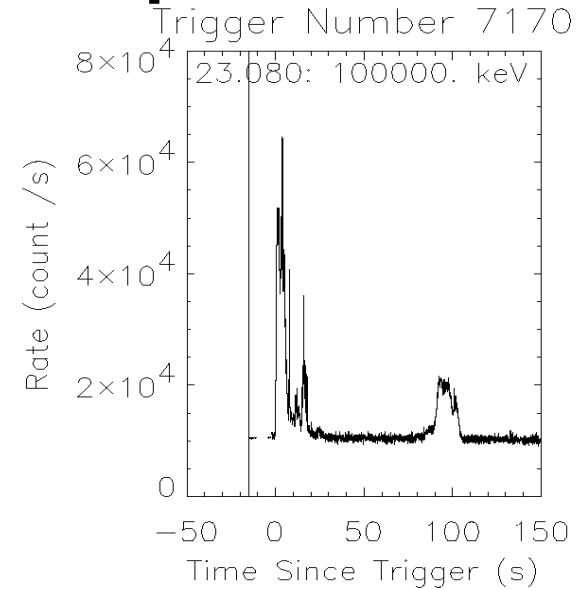
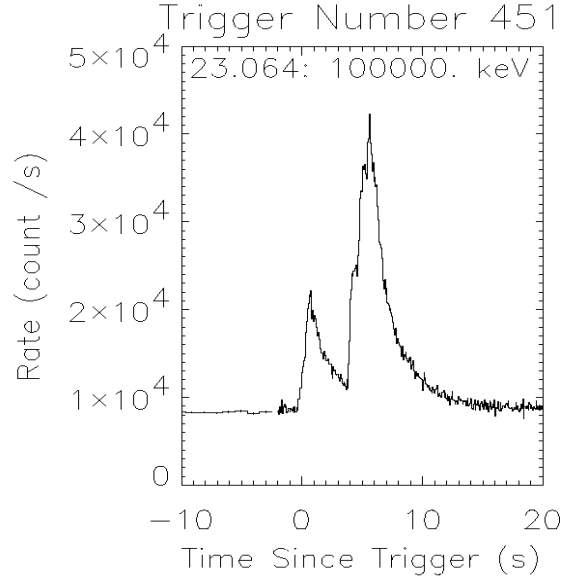
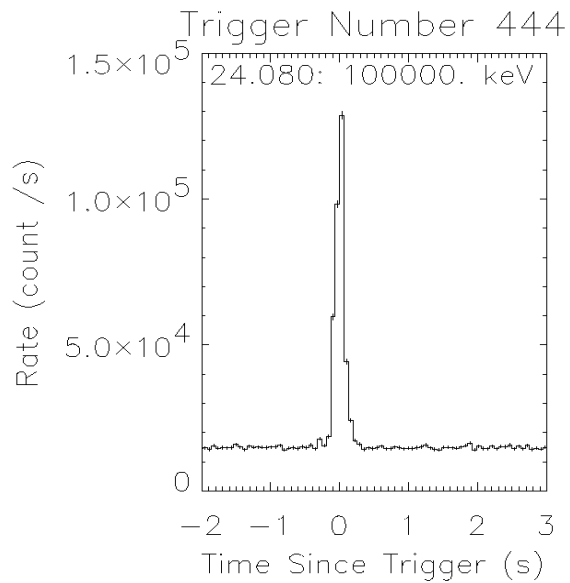
- First thoughts on GRB Observations:
 - Good field (nothing is known & high S/N)
- Need Large area to get many; remembered $-3/2$ power law: will give ~several GRBs/day
- Low-cost balloon flights others had same idea
- - propose for GRO (>>after Ltr. of Intent was due!

- Low cost; propose as signal to others.
- Propose dodecahedron (12) - 6 on top & 6 on bottom)
- Bottom was accepted; not the top, with 2 spares
- Fly the spares; need full sky w. same detectors; eight det. only (later a spare was approved

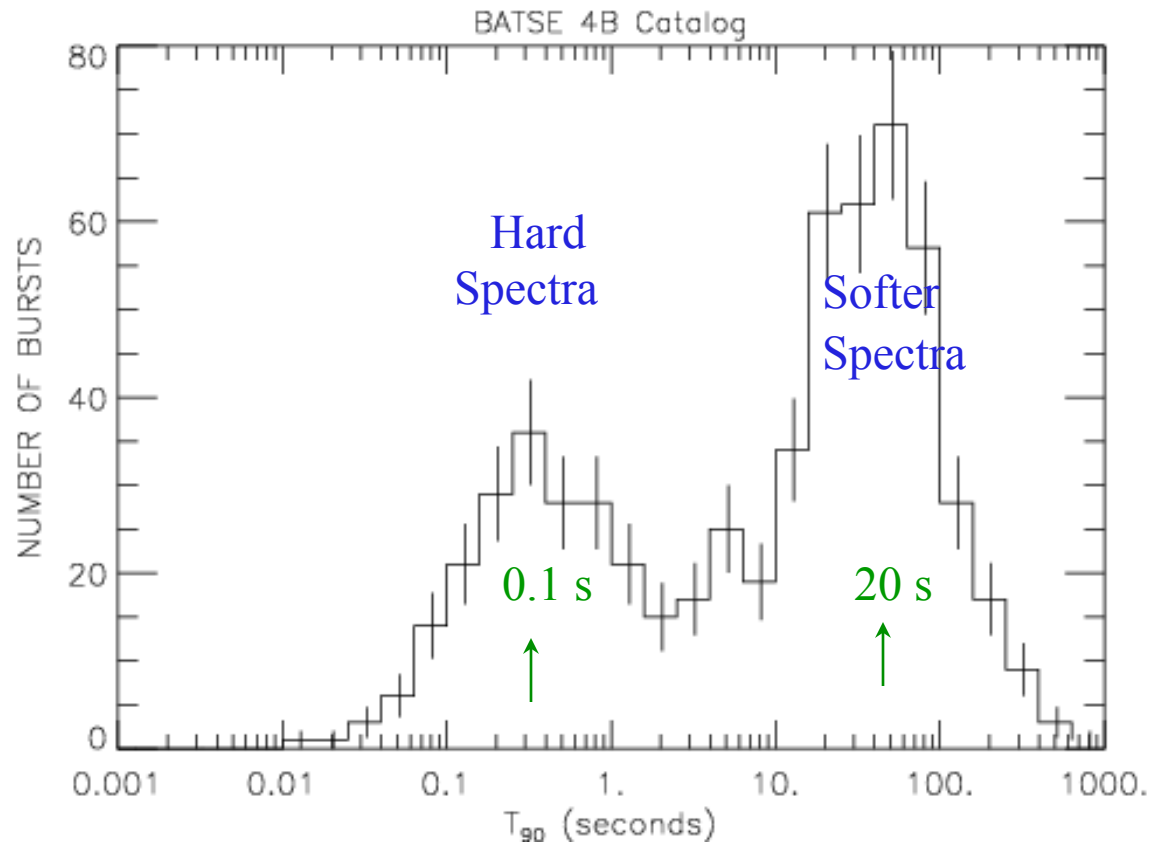
A Few Major BATSE GRB Results

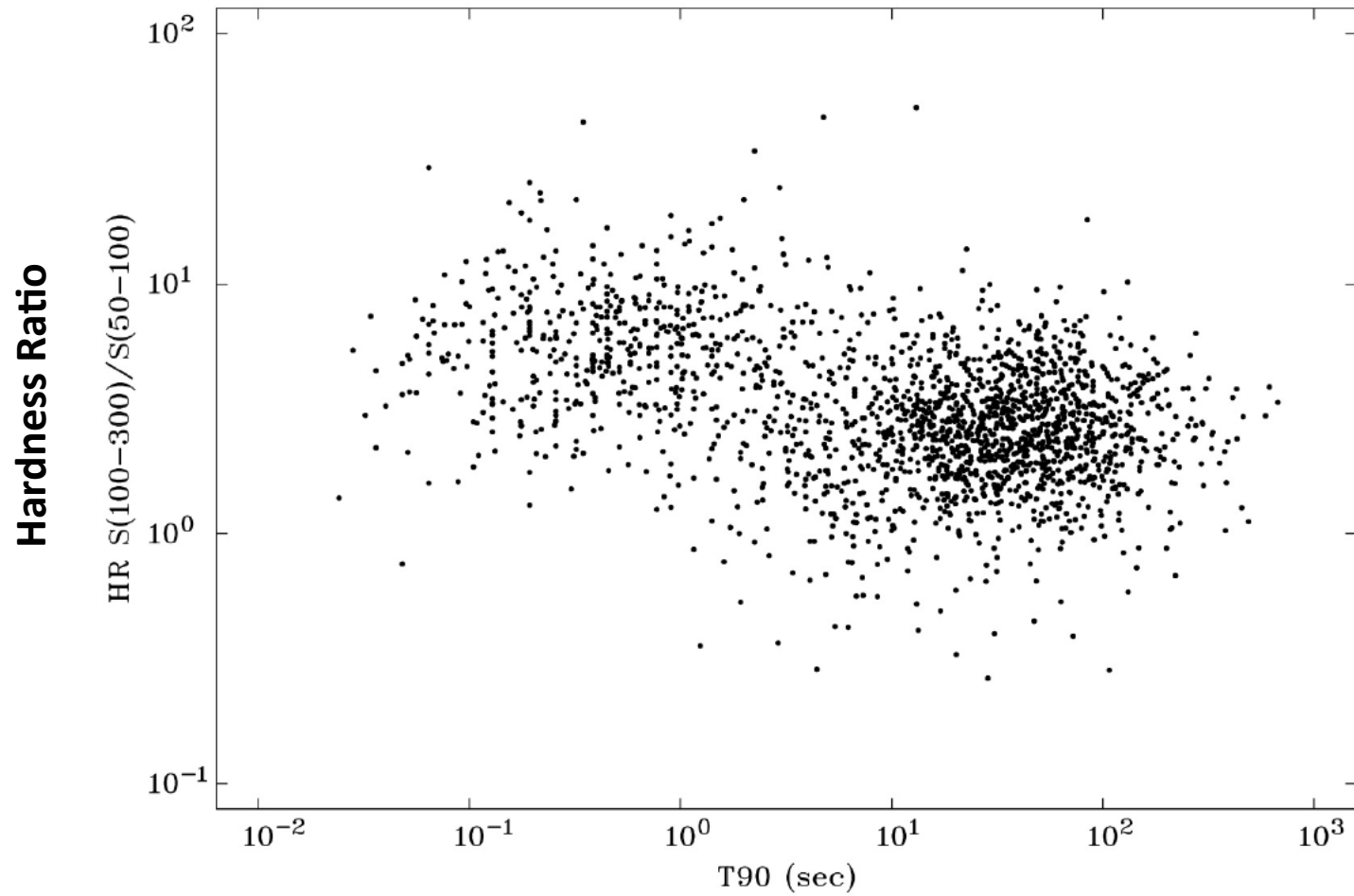
1991-2000

Diversity of GRB Profiles & Coupled Spectral / Temporal Properties



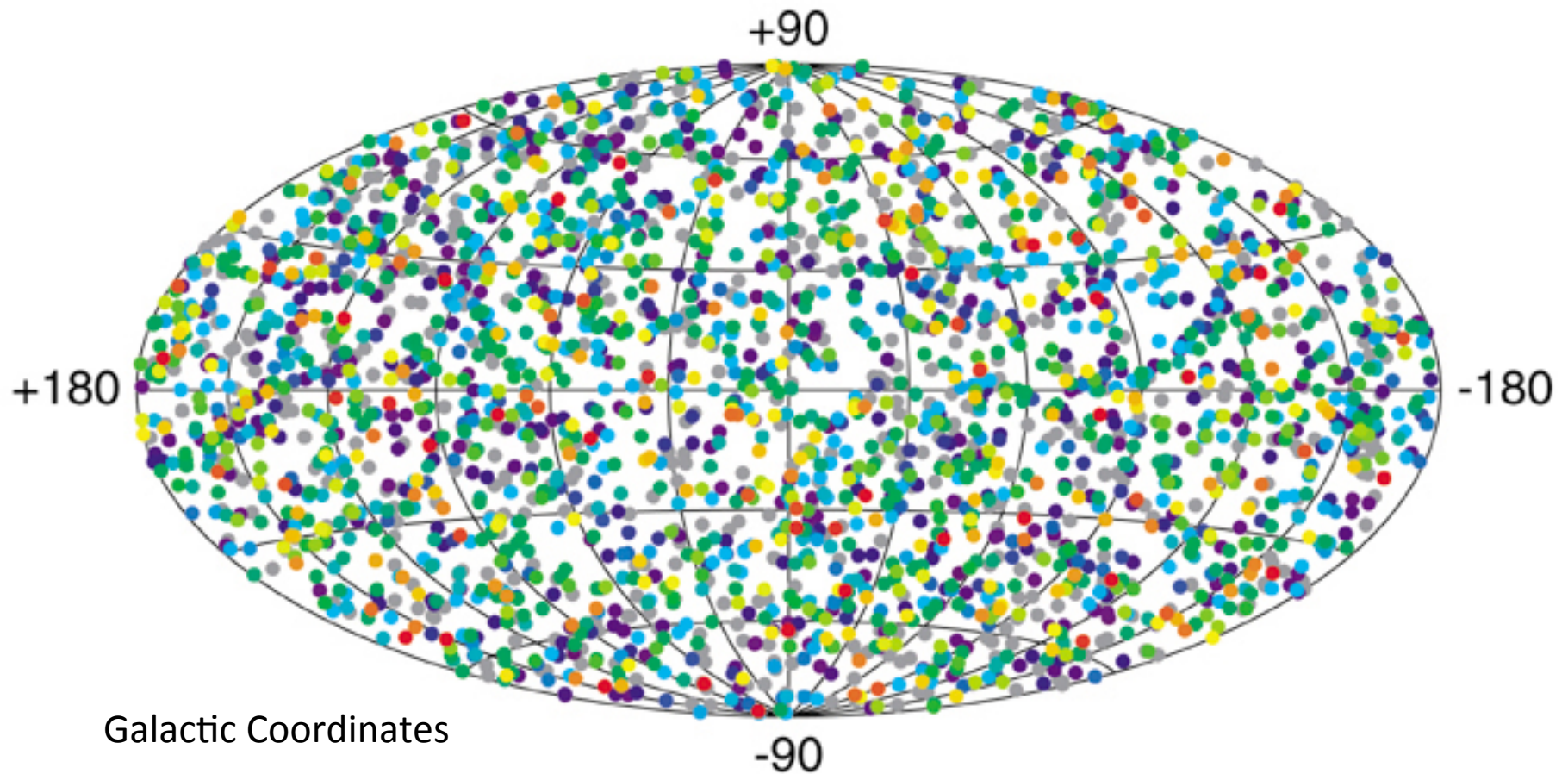
Two Distinct subclasses of γ -ray bursts: short/hard & long/soft



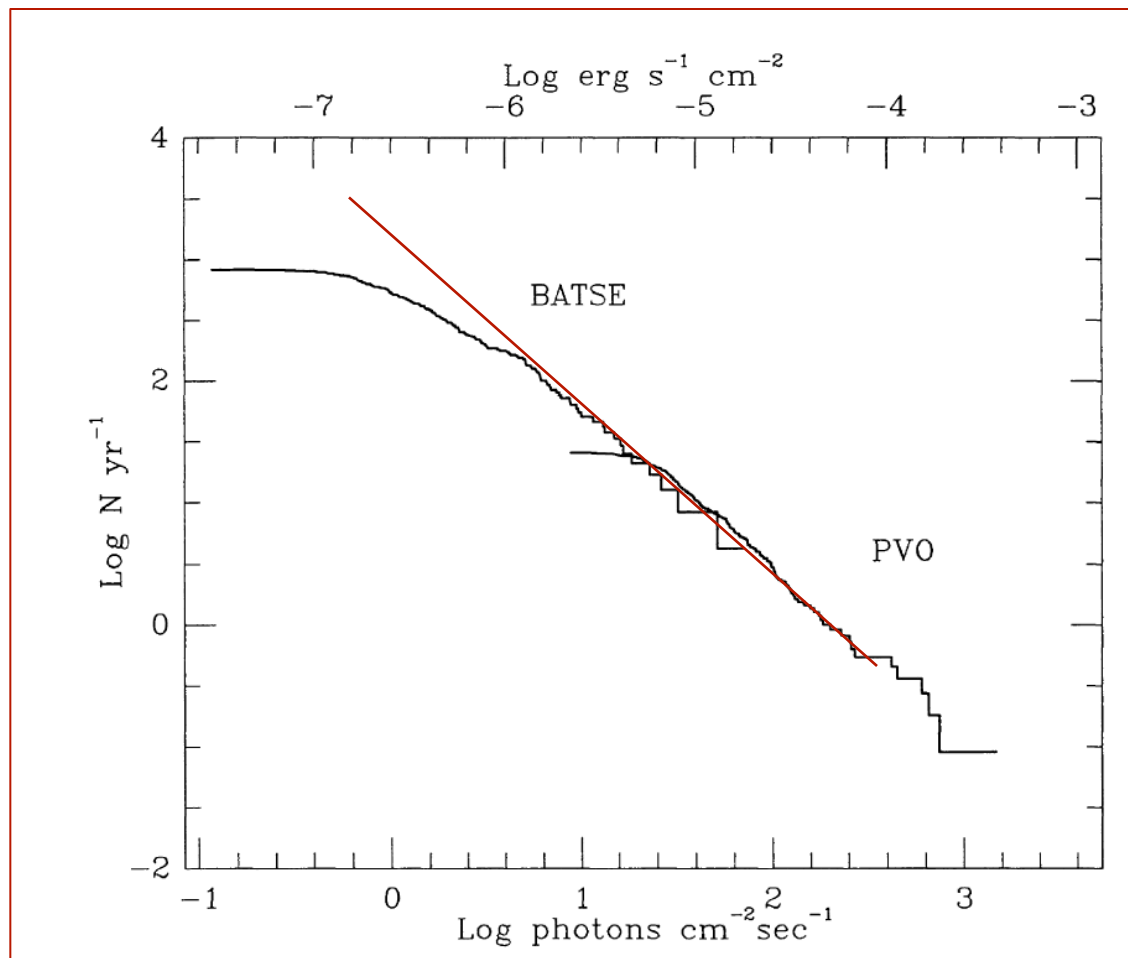


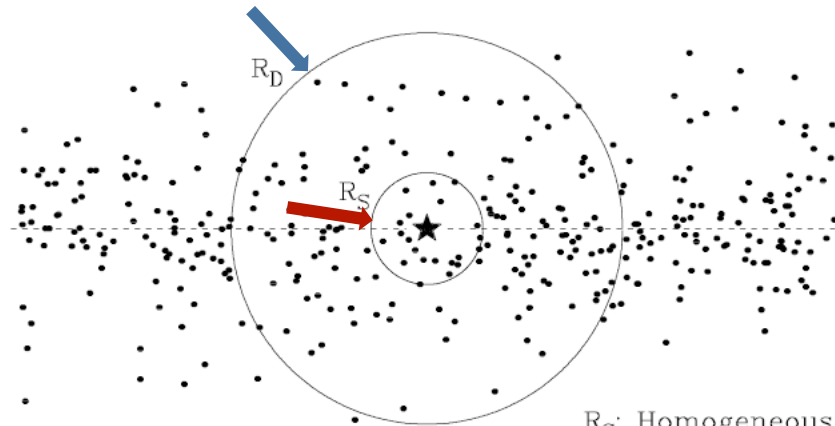
Duration of Gamma-ray Bursts (sec)

2704 BATSE Gamma-Ray Bursts

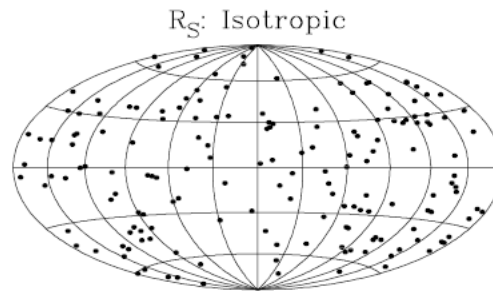


Apr. 1991 – May 2000

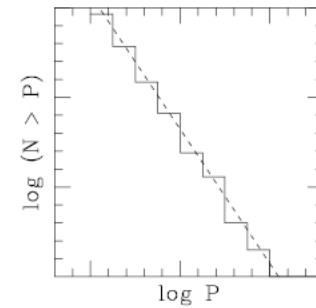




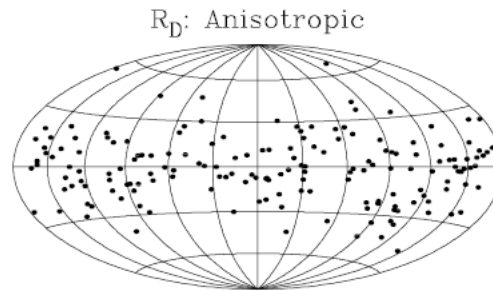
R_S - Galactic, Nearby



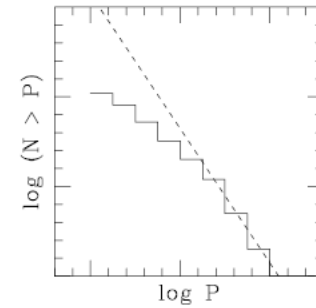
R_S : Homogeneous

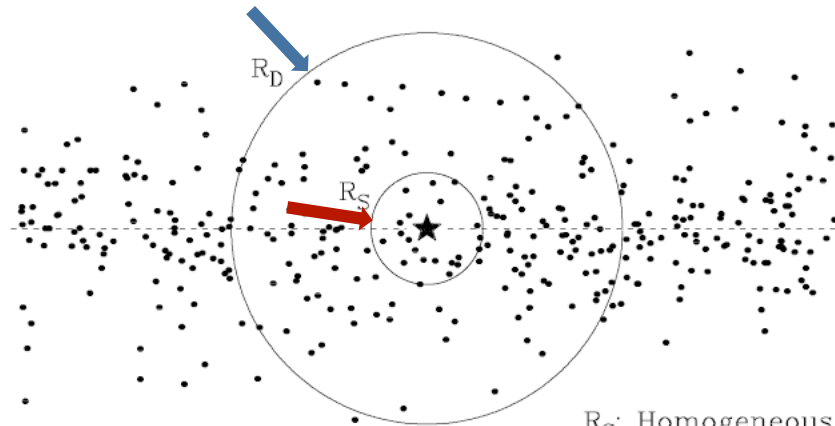


R_D - Galactic, Far away



R_D : Inhomogeneous

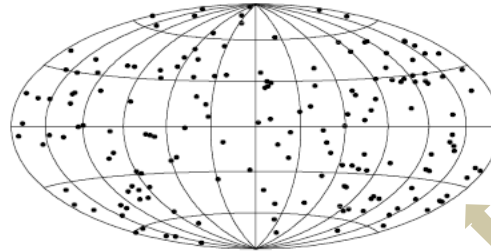




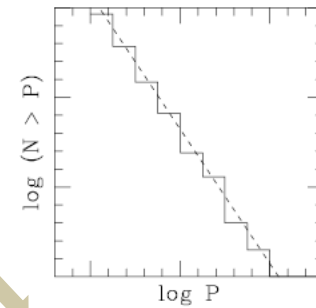
R_S - Galactic, Nearby



R_S : Isotropic



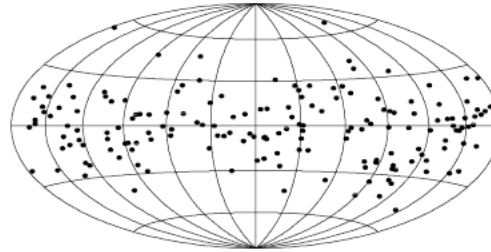
R_S : Homogeneous



R_D - Galactic, Far away



R_D : Anisotropic



R_D : Inhomogeneous

