

Short Gamma-Ray Burst Afterglows

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Munich, Germany , 05.08.2012

special thanks to:

Raffaella Margutti

B.Ashley Zauderer

Ian Czekala

Ryan Chornock

Eleonora Troja

Takanori Sakamoto

Neil Gehrels

Derek Fox

Philipp Podsiadlowski

Physics of GRBs: The big picture

Progenitor

Central engine

Internal shocks

External forward shock

γ -rays

X-ray
Optical
NIR
Radio

Unanswered, but reachable questions:

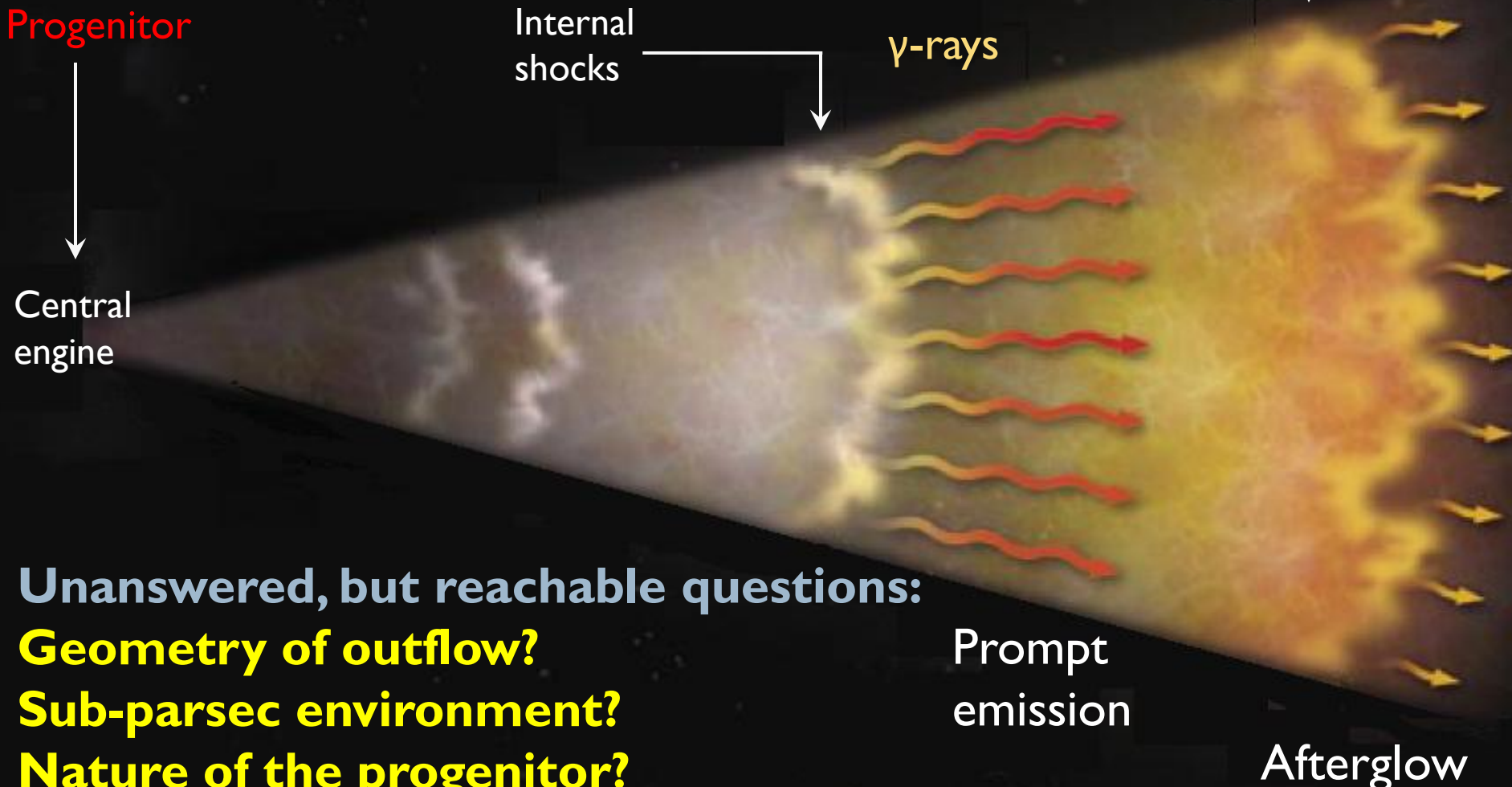
Geometry of outflow?

Sub-parsec environment?

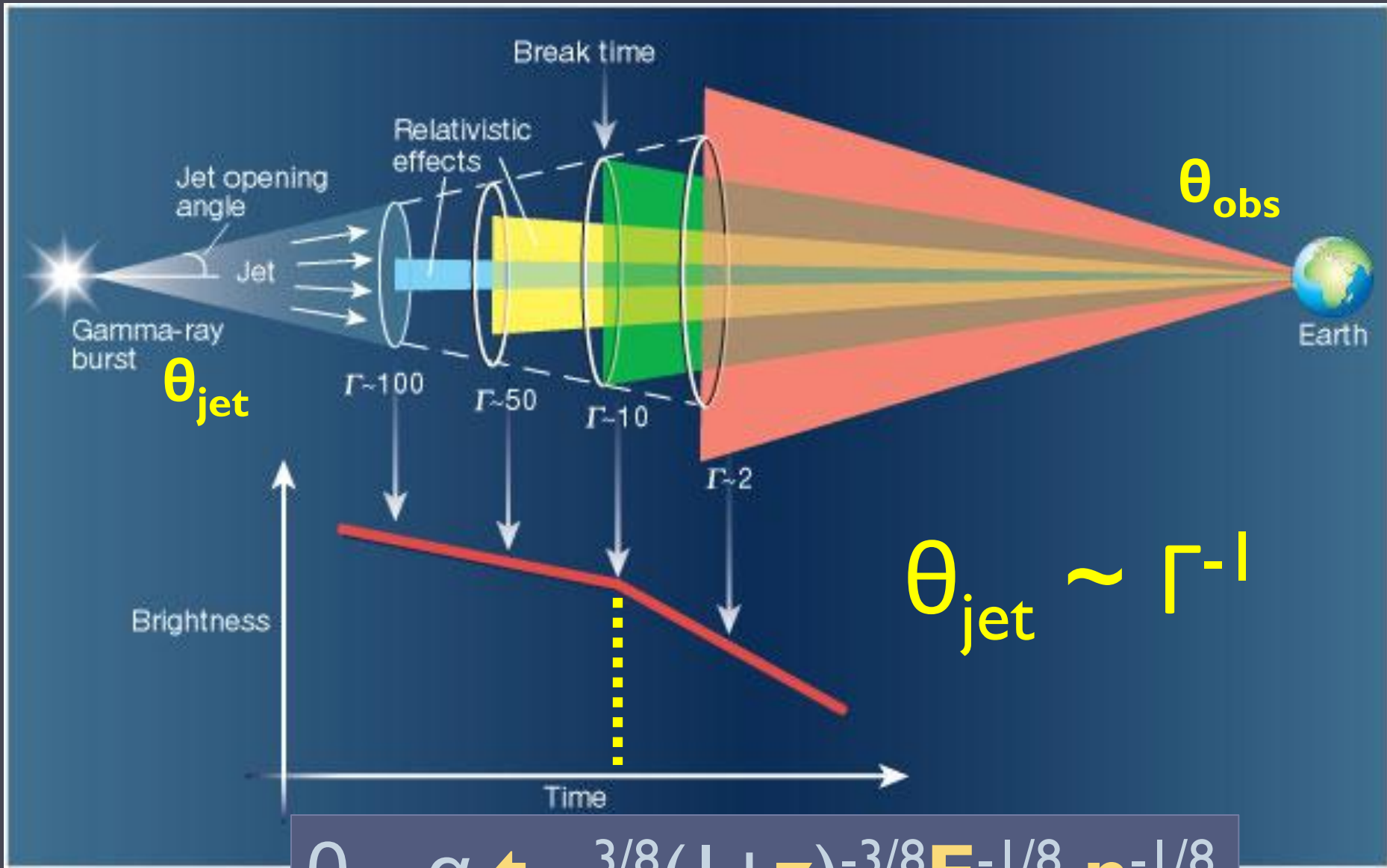
Nature of the progenitor?

Prompt emission

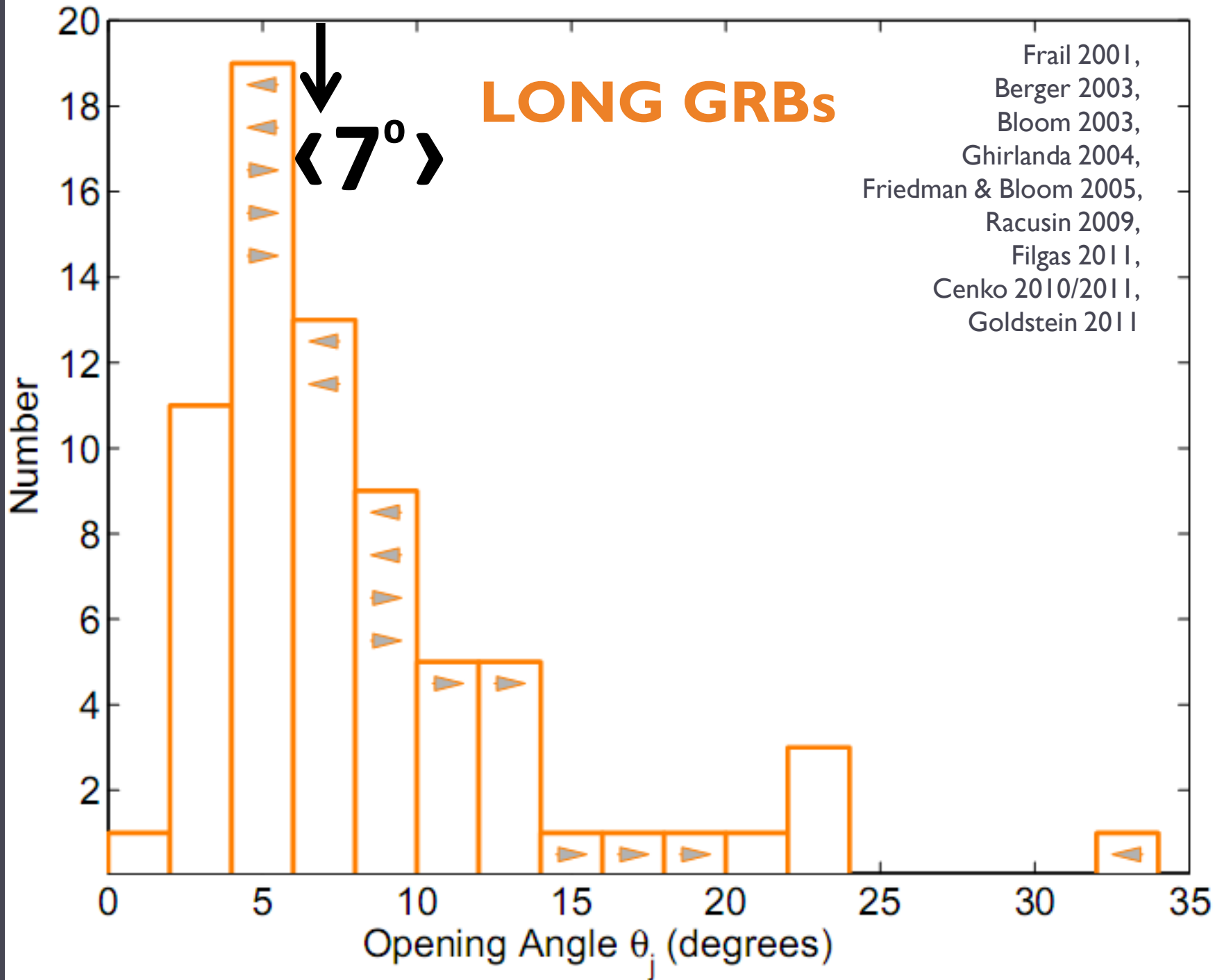
Afterglow

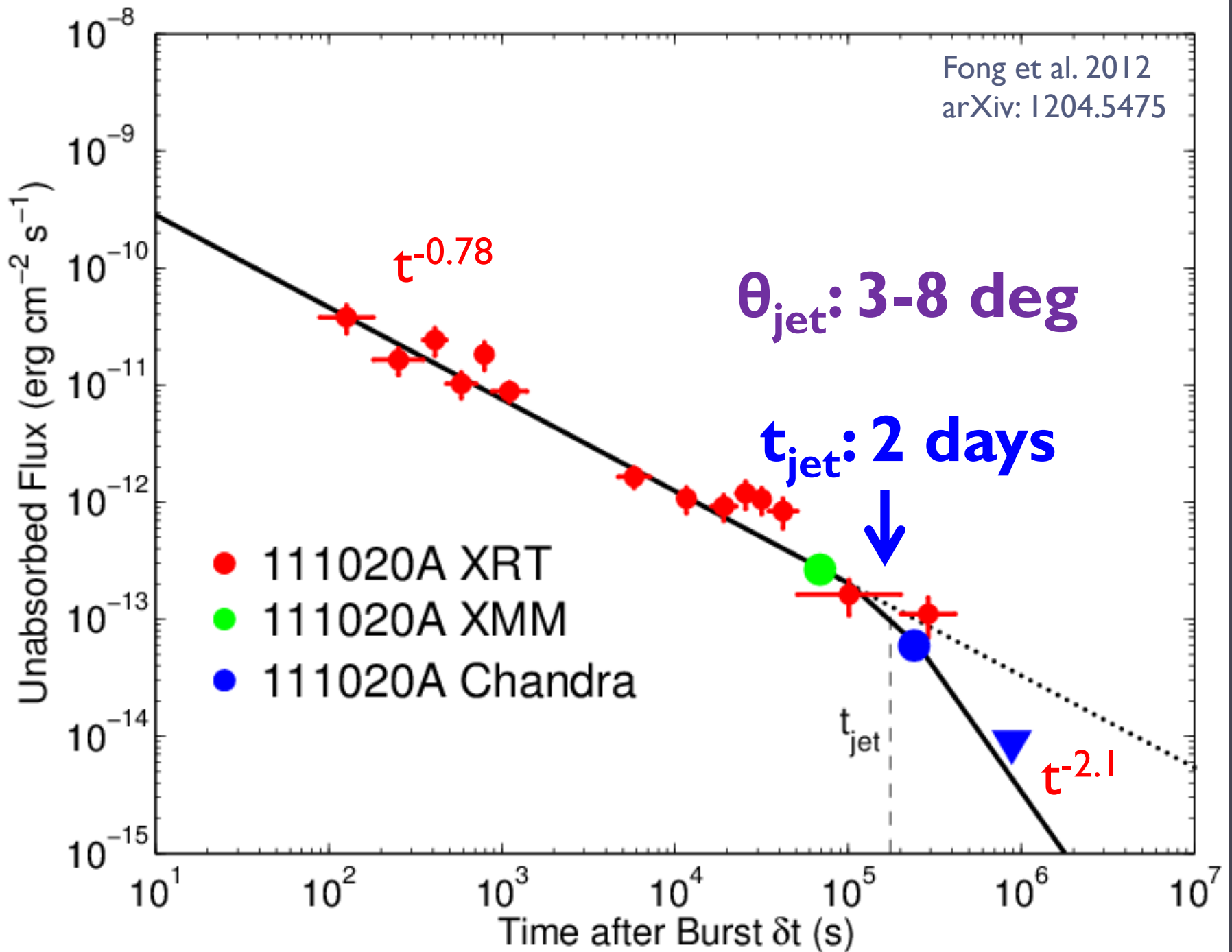


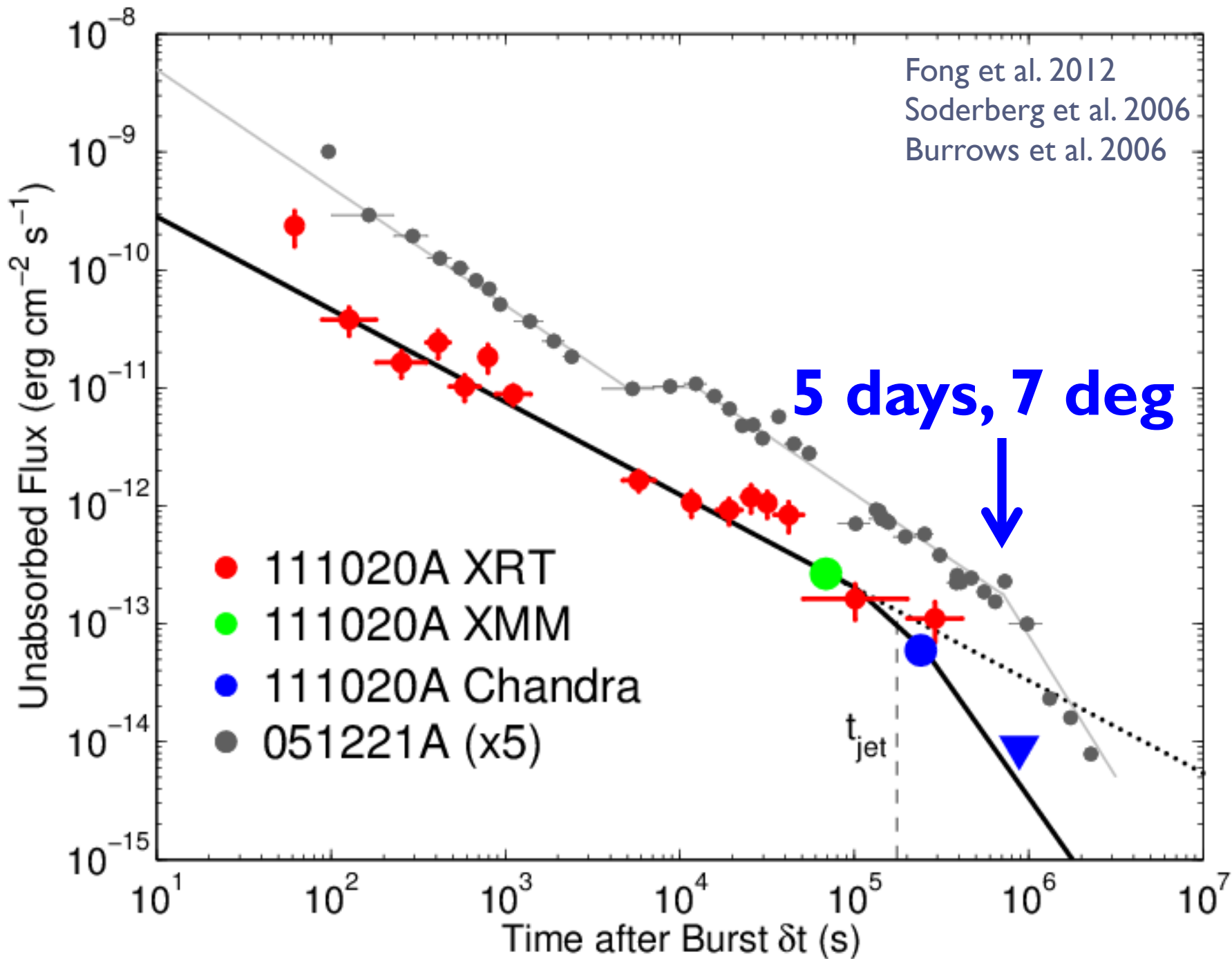
Geometry of outflow? Jet Breaks



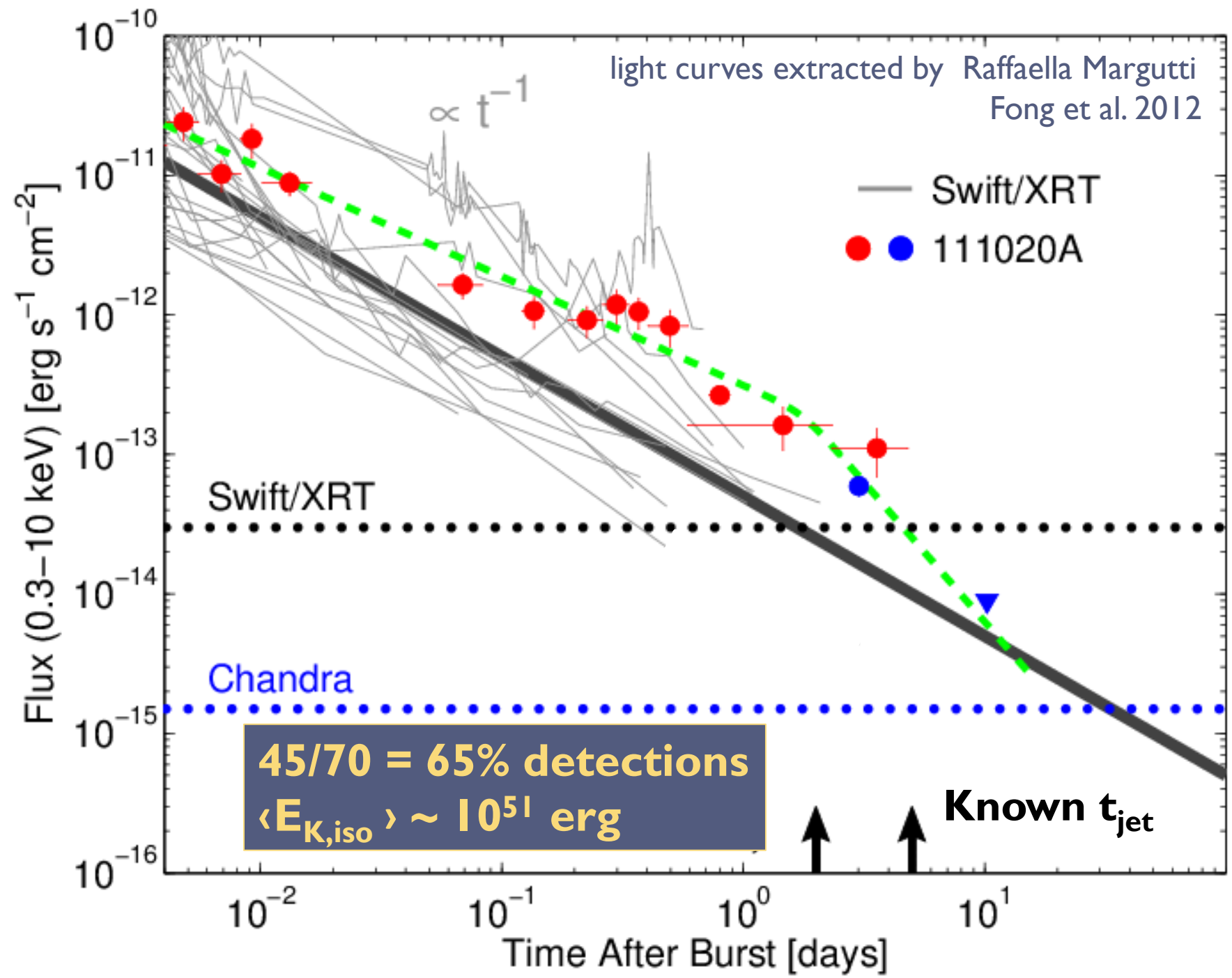
$$\theta_{jet} \propto t_{jet}^{3/8} (1+z)^{-3/8} E^{-1/8} n^{-1/8}$$





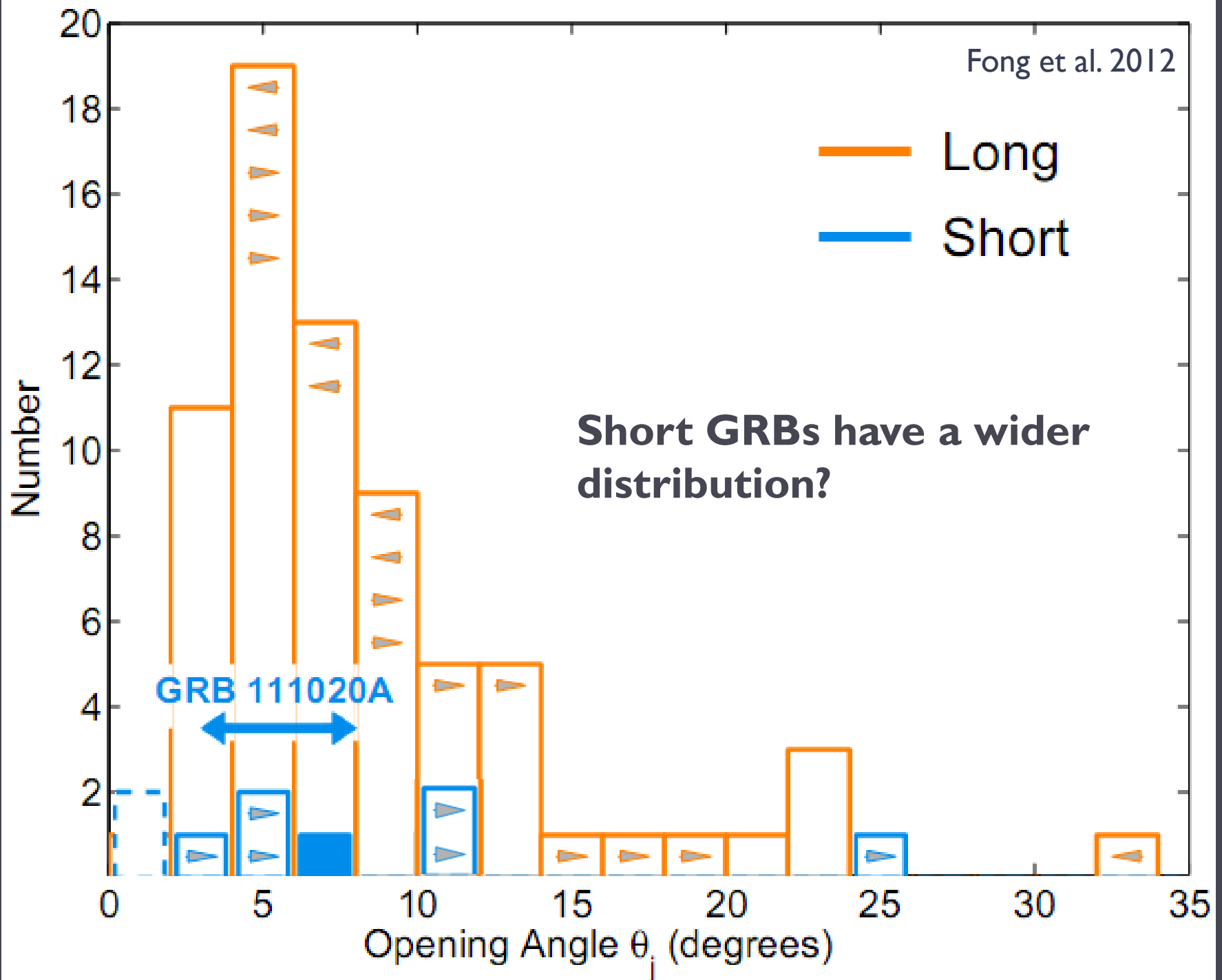


light curves extracted by Raffaella Margutti
Fong et al. 2012



45/70 = 65% detections
⟨E_{K,iso}⟩ ~ 10⁵¹ erg

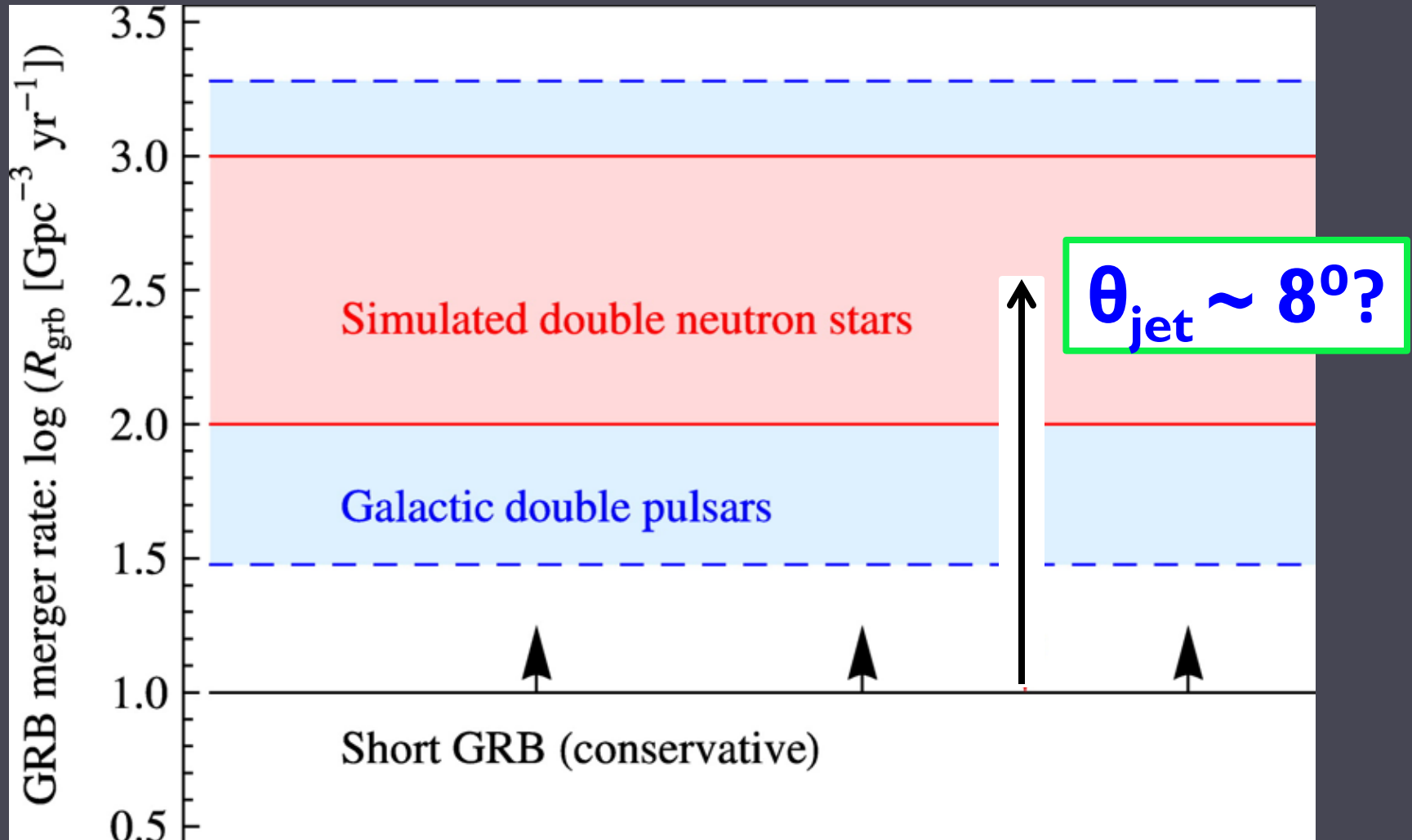
Known t_{jet}



Implications: Energy Scale

- ⊙ isotropic-equivalent energy $E_{\text{iso}} \sim 10^{51}$ erg
 - ⊙ beaming-corrected true energies $\sim 10^{49}$ erg
 - ⊙ mechanism of energy extraction predictions:
 - neutrino-antineutrino annihilation $\sim 10^{48}$ - 10^{49} erg
 - MHD processes, magnetically-dom. jet $\sim >10^{49}$ erg
- Rosswog et al. 2003, Rosswog 2005, Birkel et al. 2007, Lee & Ramirez-Ruiz 2007

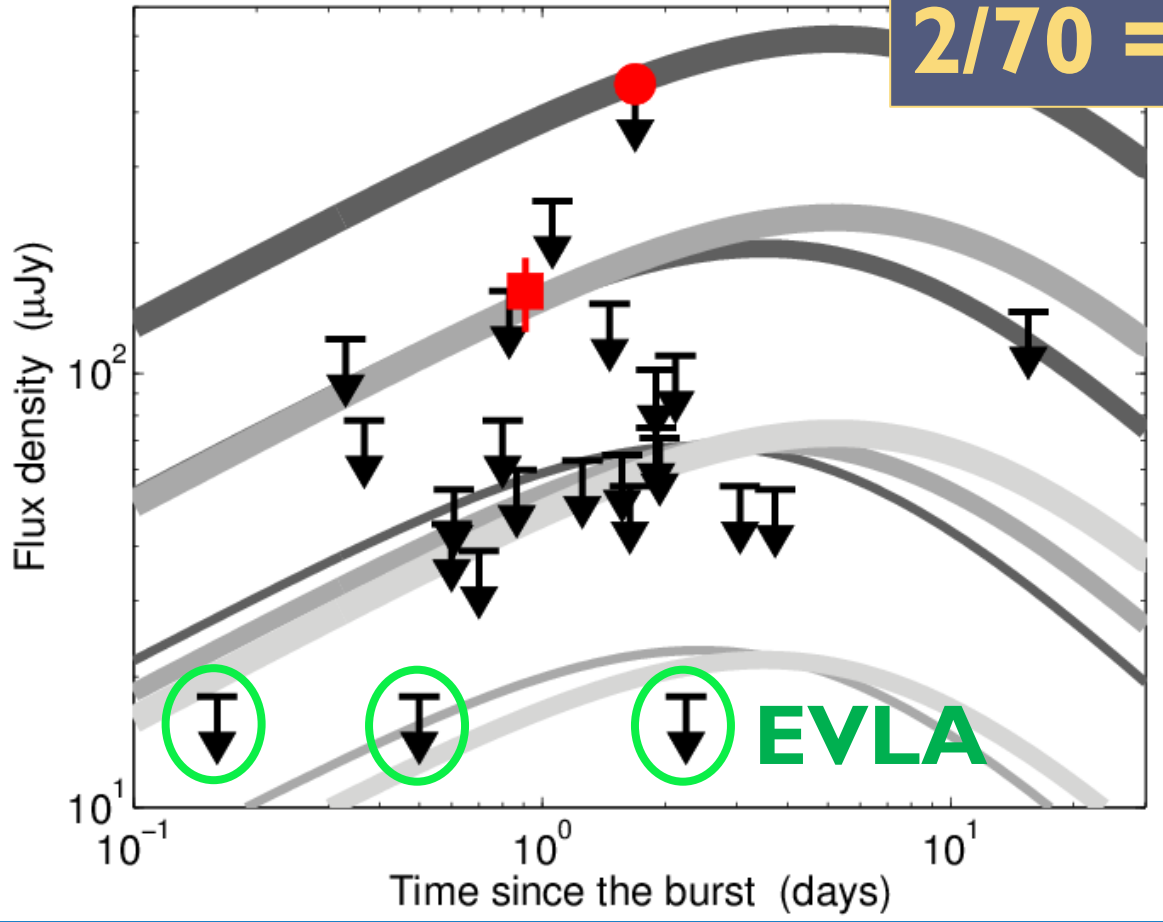
Implications: Rates



observed short GRB rate of $10 \text{ Gpc}^{-3} \text{ yr}^{-1} \rightarrow 100\text{-}1000 \text{ Gpc}^{-3} \text{ yr}^{-1}$

Sub-pc environment: Radio afterglows

2/70 = 3% detections

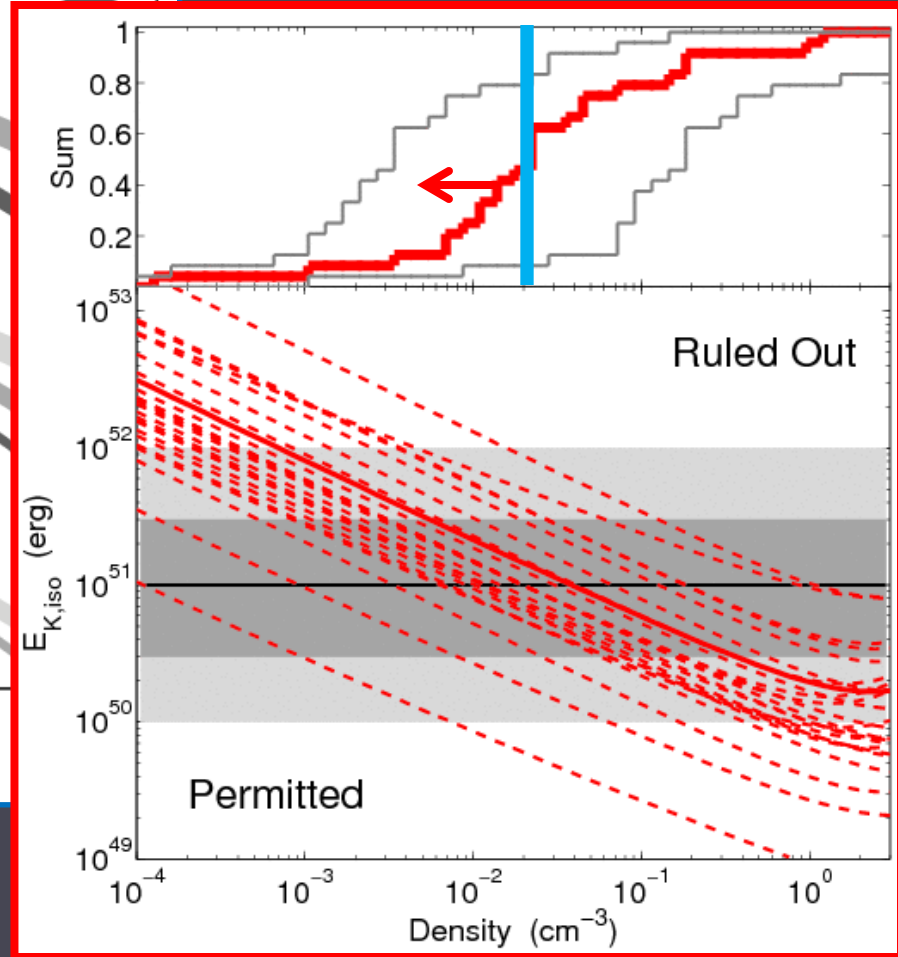
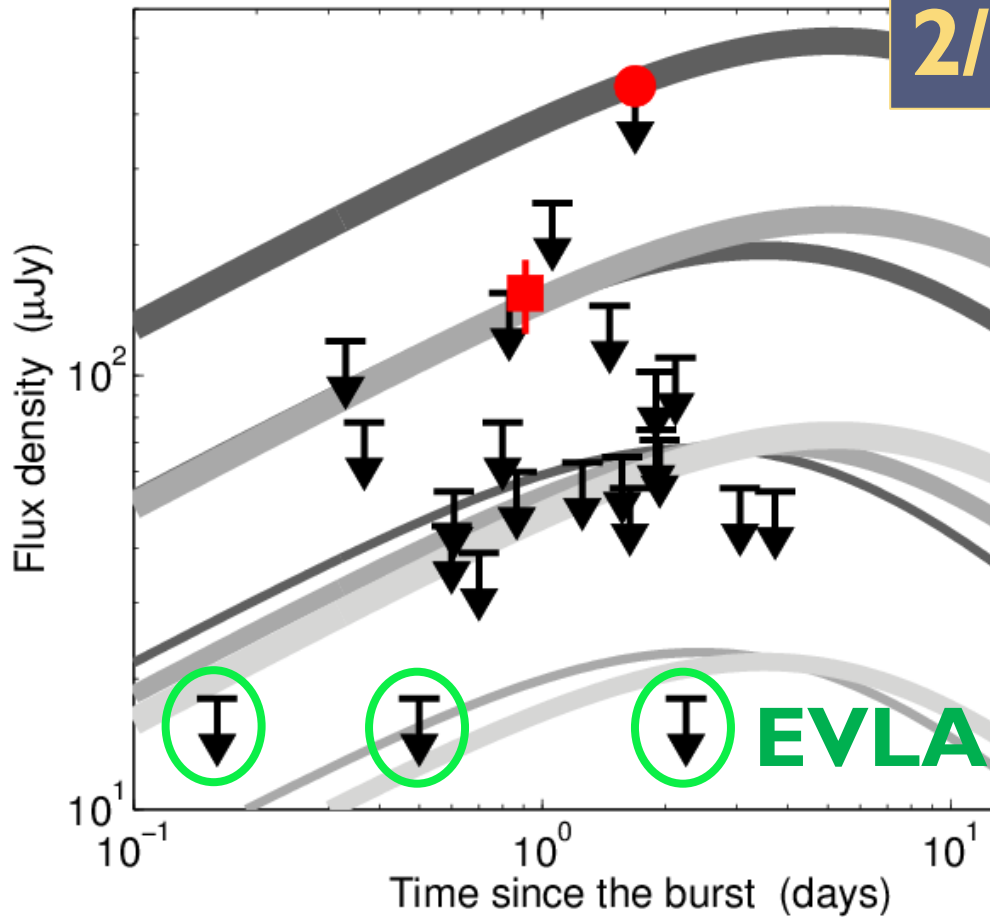


Fong in prep. 2012

$$F_{radio,\nu} \propto n^{1/2} E_{K,iso}^{5/6}$$

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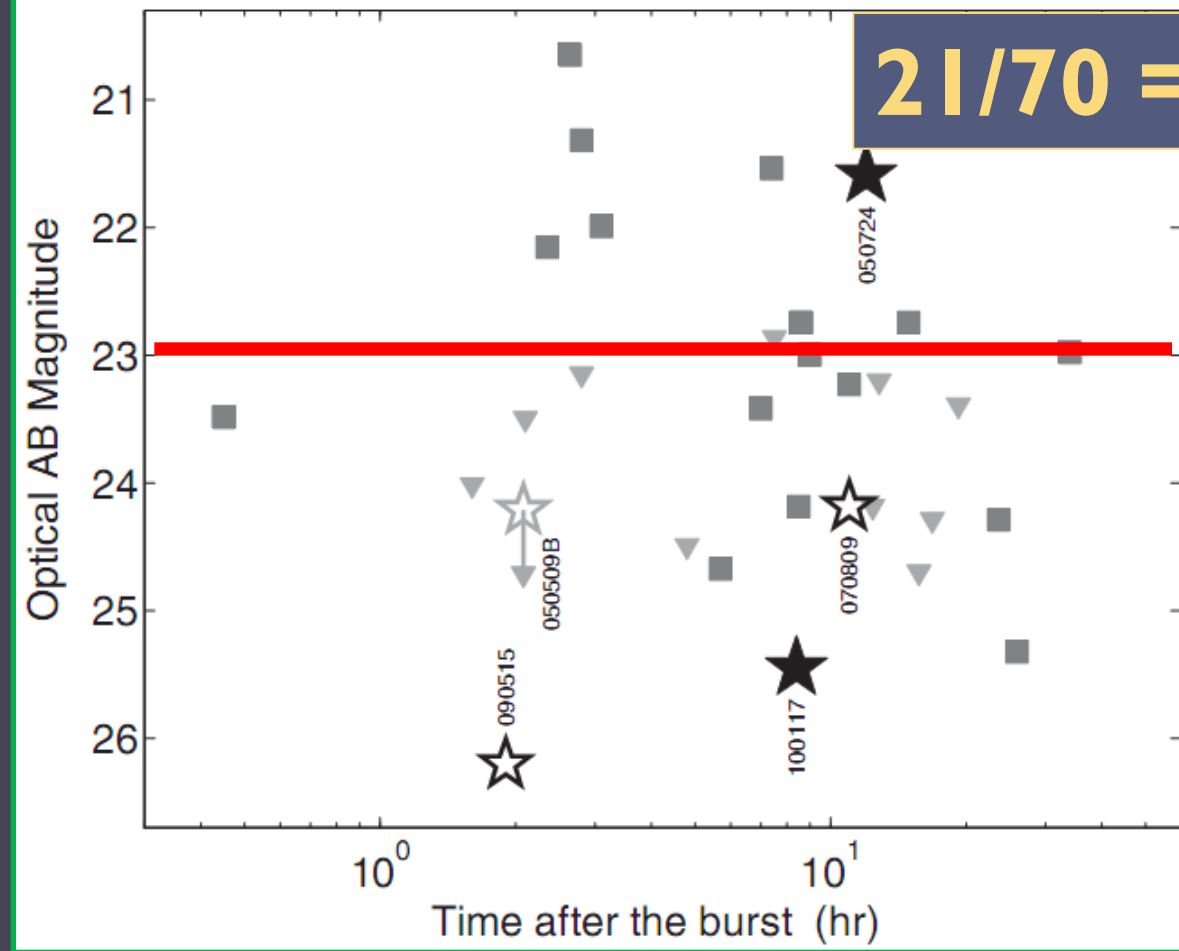
Fong in prep. 2012

$$F_{\text{radio},\nu} \propto n^{1/2} E_{K,\text{iso}}^{5/6}$$

from UL and detections: $\langle n \rangle \leq 0.02 \text{ cm}^{-3}$

Sub-pc environment: Optical afterglows

21/70 = 30% detections

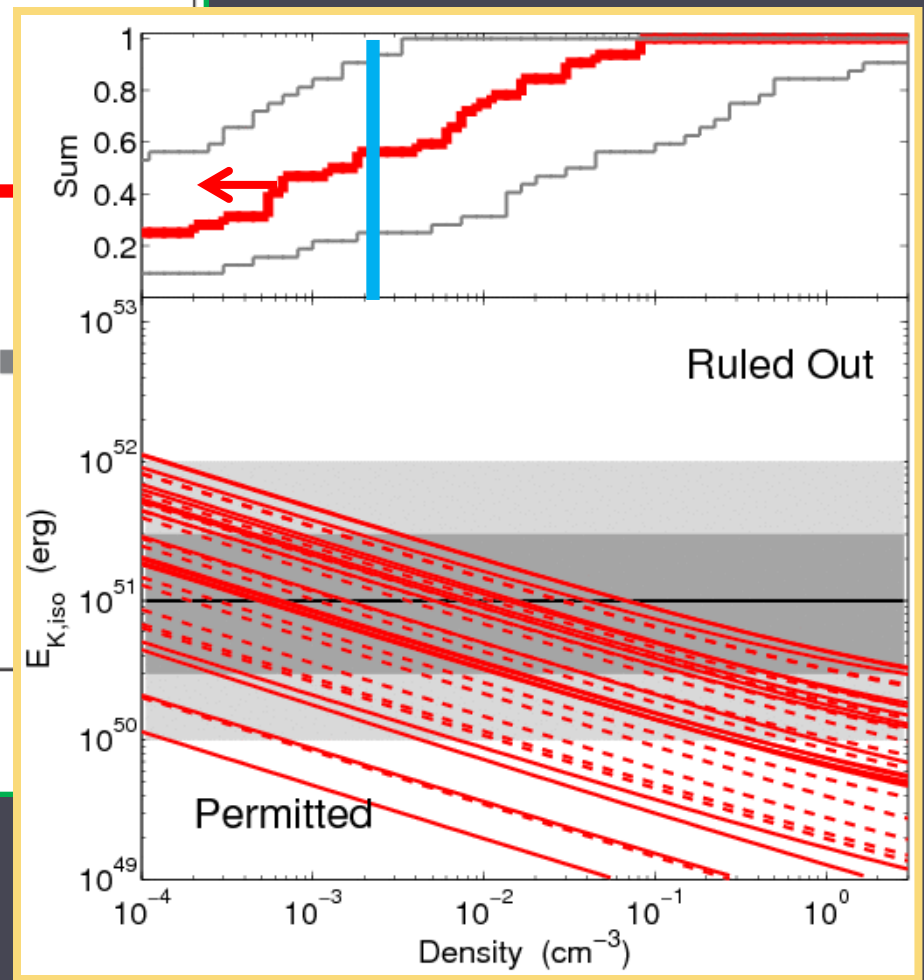
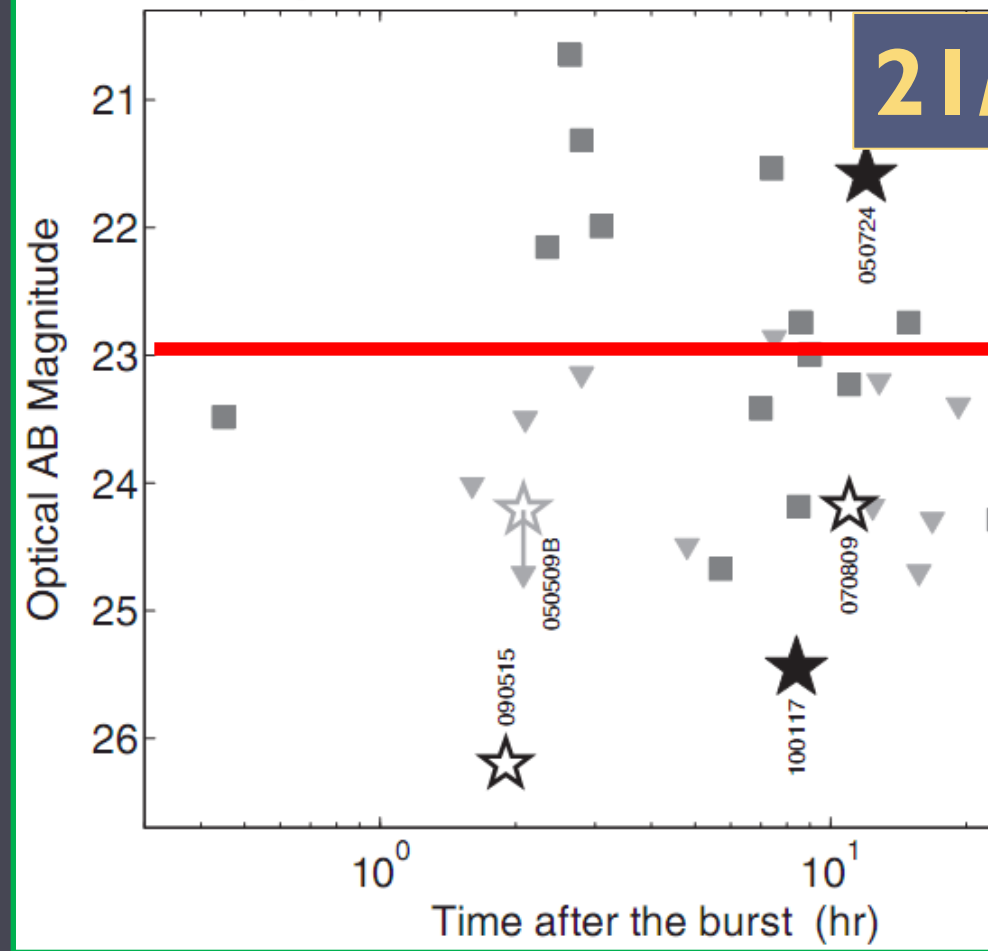


Berger 2010, Fong et al. 2011

$$F_{opt,\nu} \propto n^{1/2} E_{K,iso}^{(3+p)/4}$$

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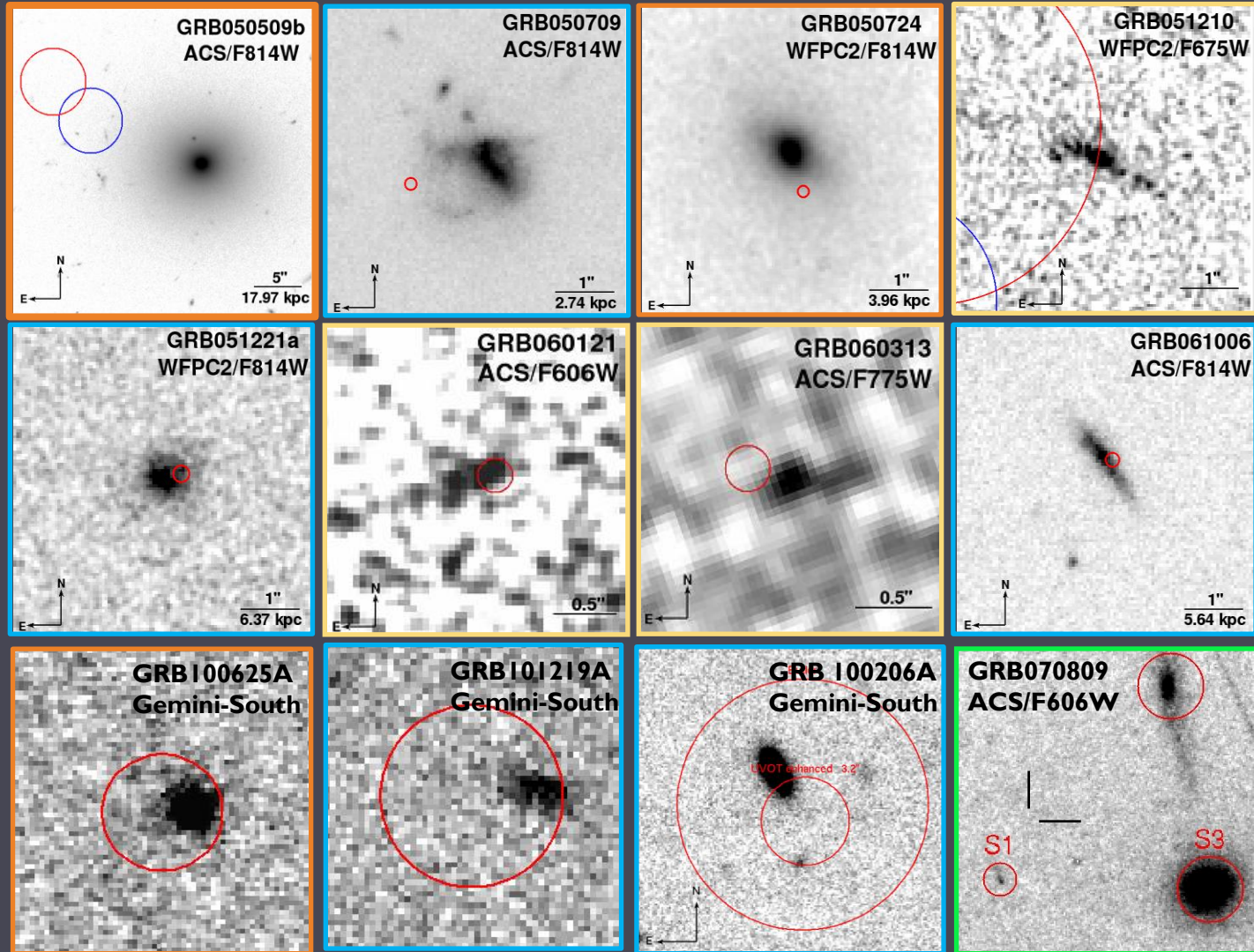


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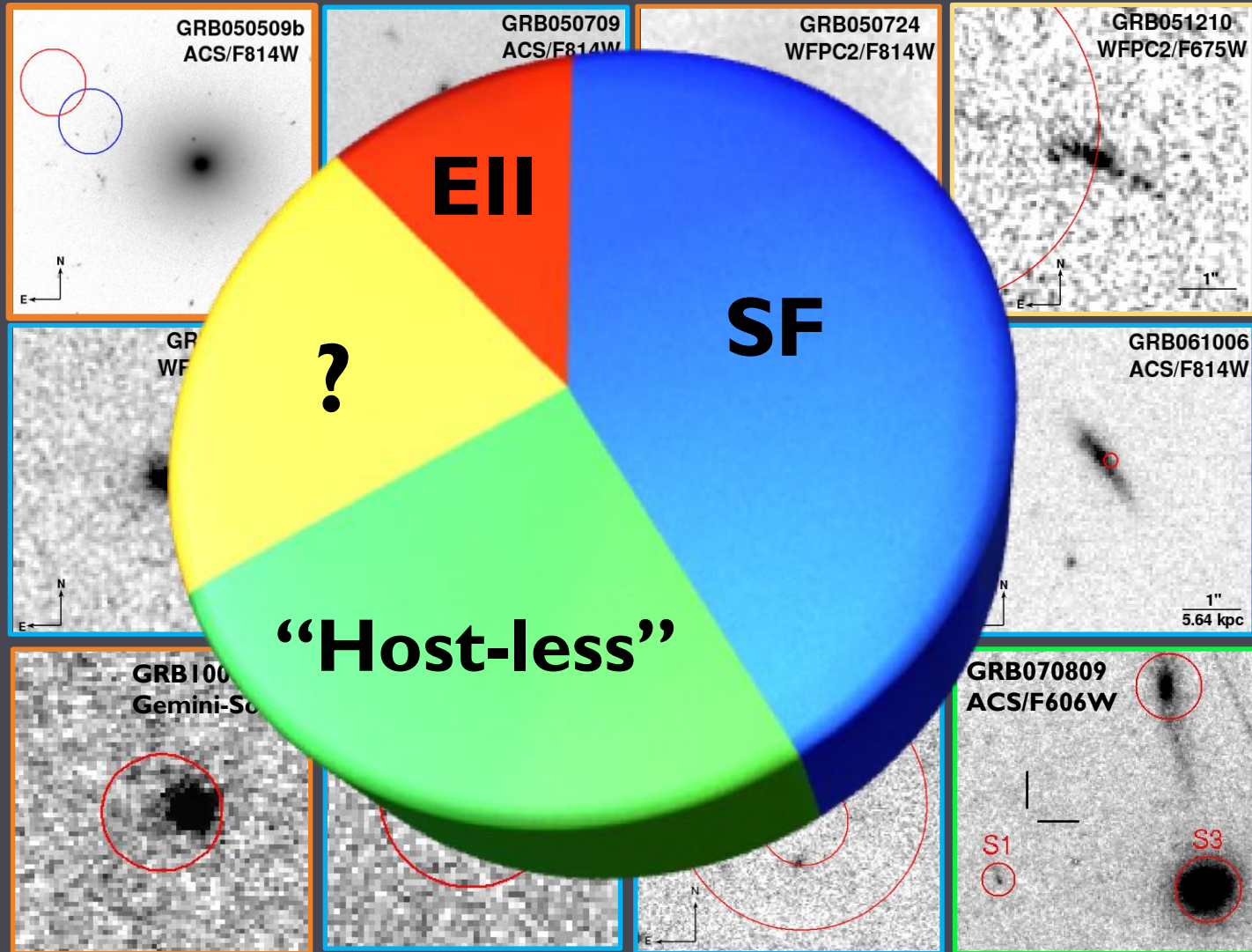
from UL and detections: $\langle n \rangle \leq 0.002 \text{ cm}^{-3}$

Nature of the progenitor? Galactic-scale environments



See posters:
D. Perley,
T.Sakamoto

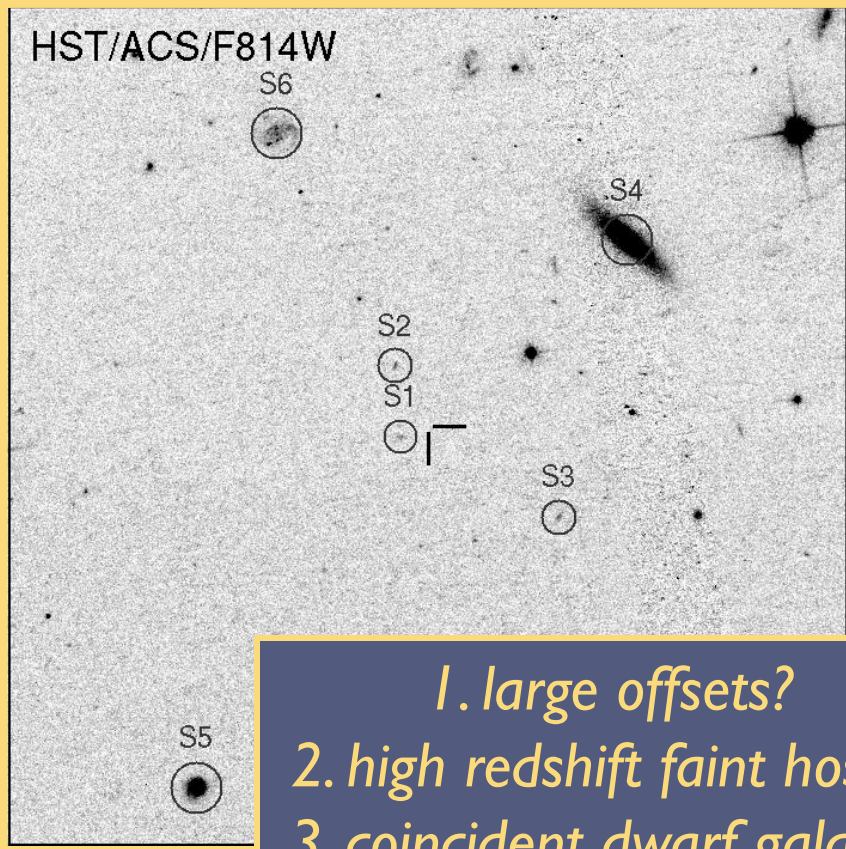
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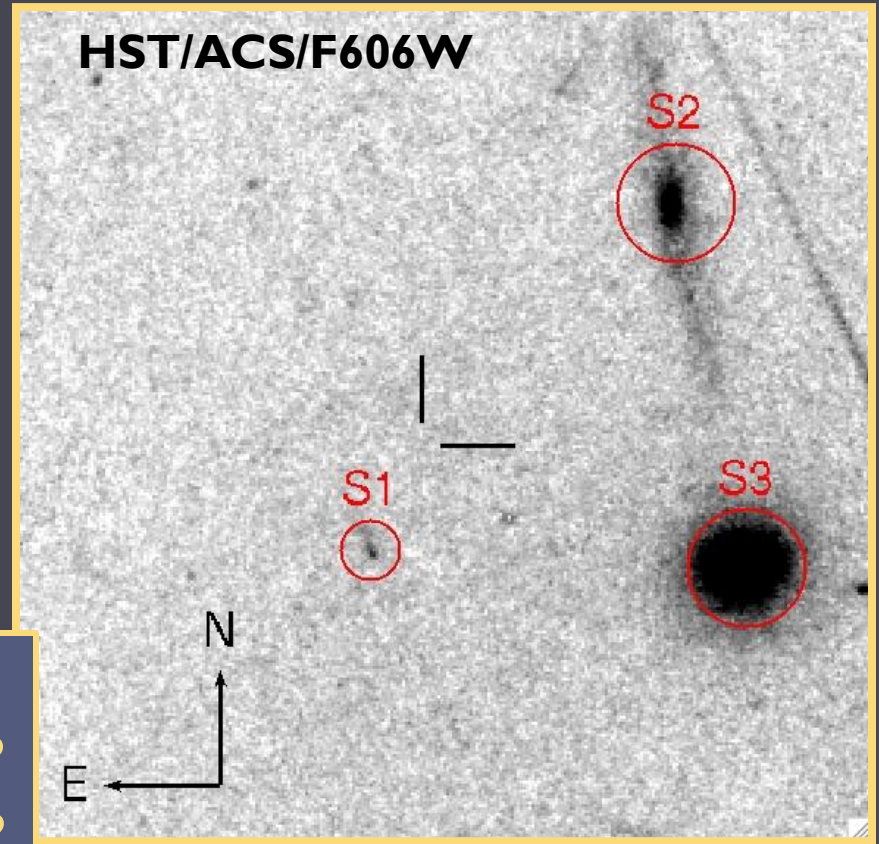
See posters:
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Evidence for kicks? Six “host-less” bursts

GRB 061201 Berger 2010; Stratta et al. 2006; Fong et al. 2010



GRB 070809

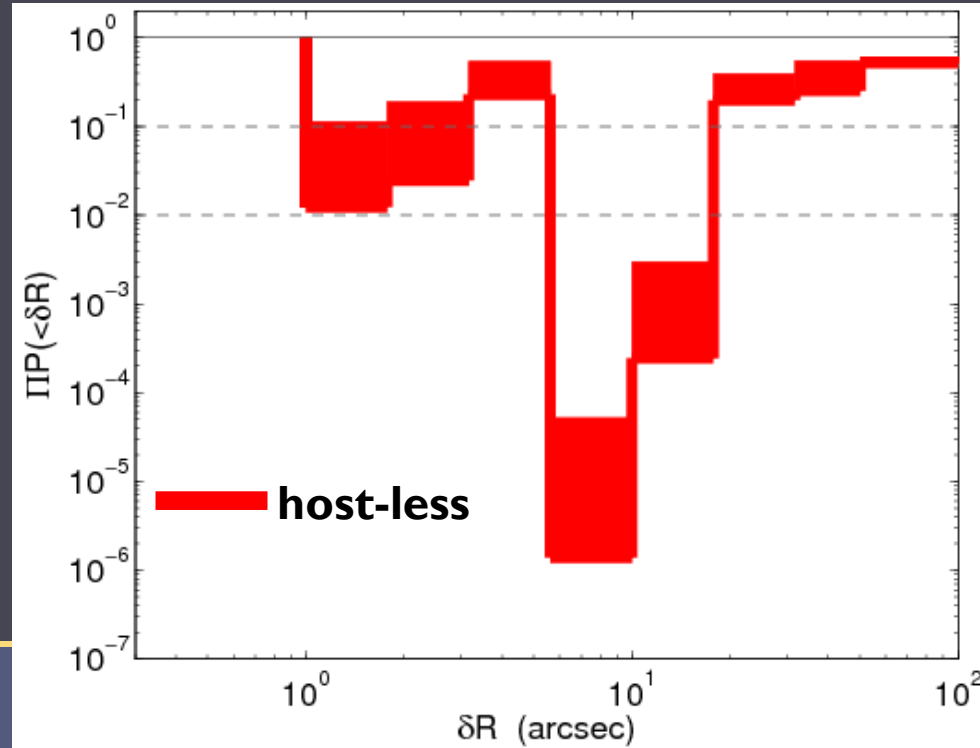
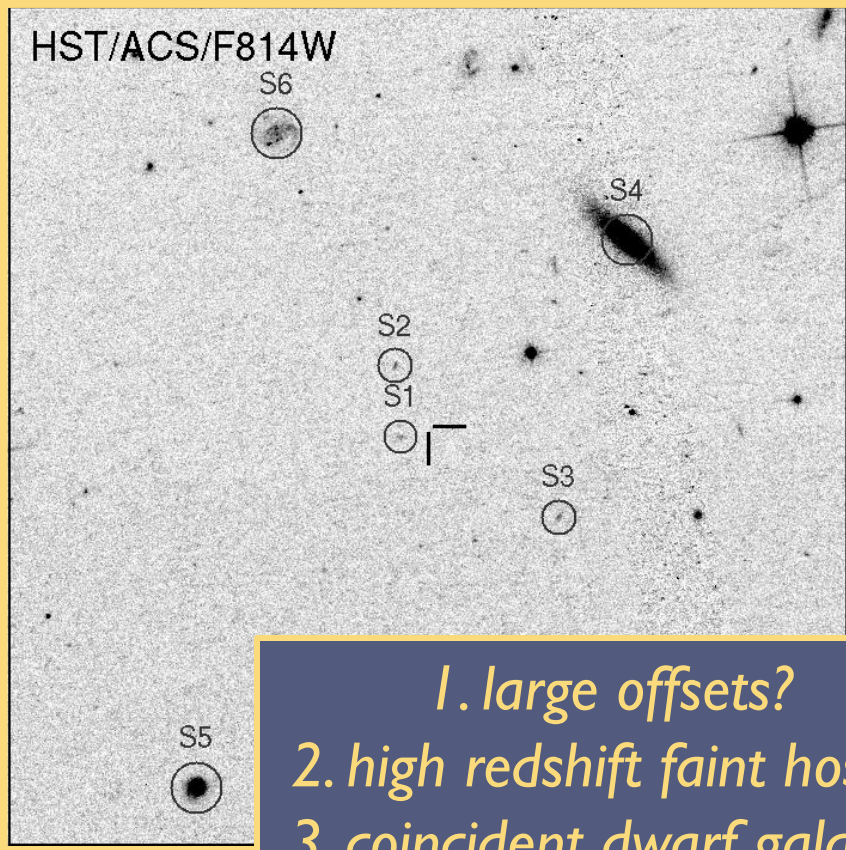


- 1. large offsets?*
- 2. high redshift faint hosts?*
- 3. coincident dwarf galaxy?*

What is the probability of chance coincidence?
(What is the likelihood of finding an unrelated galaxy?)

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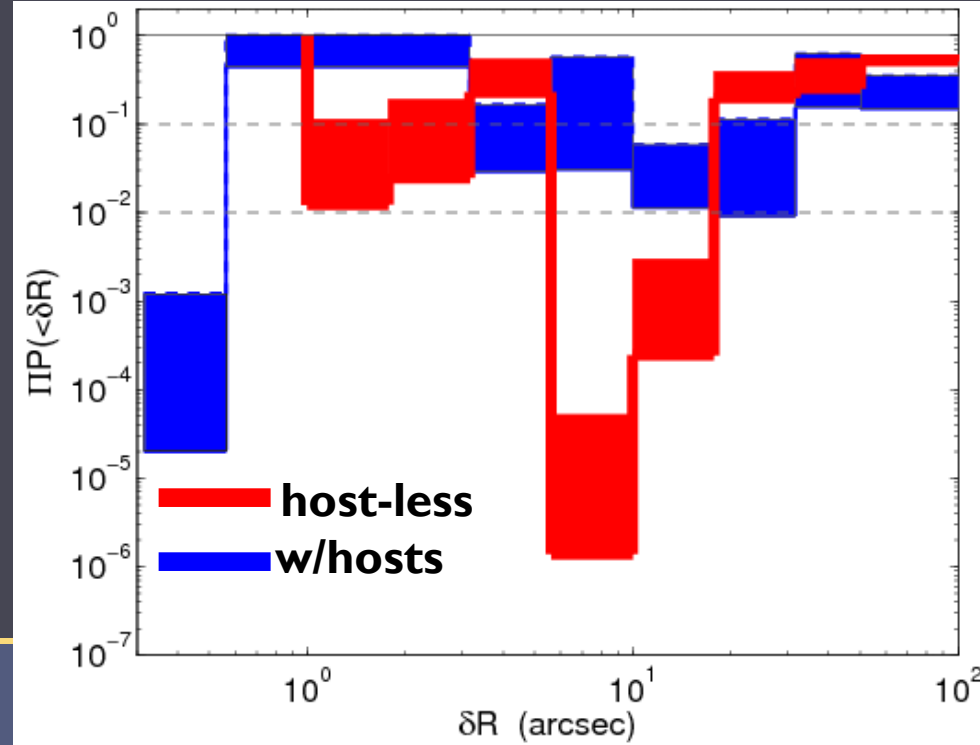
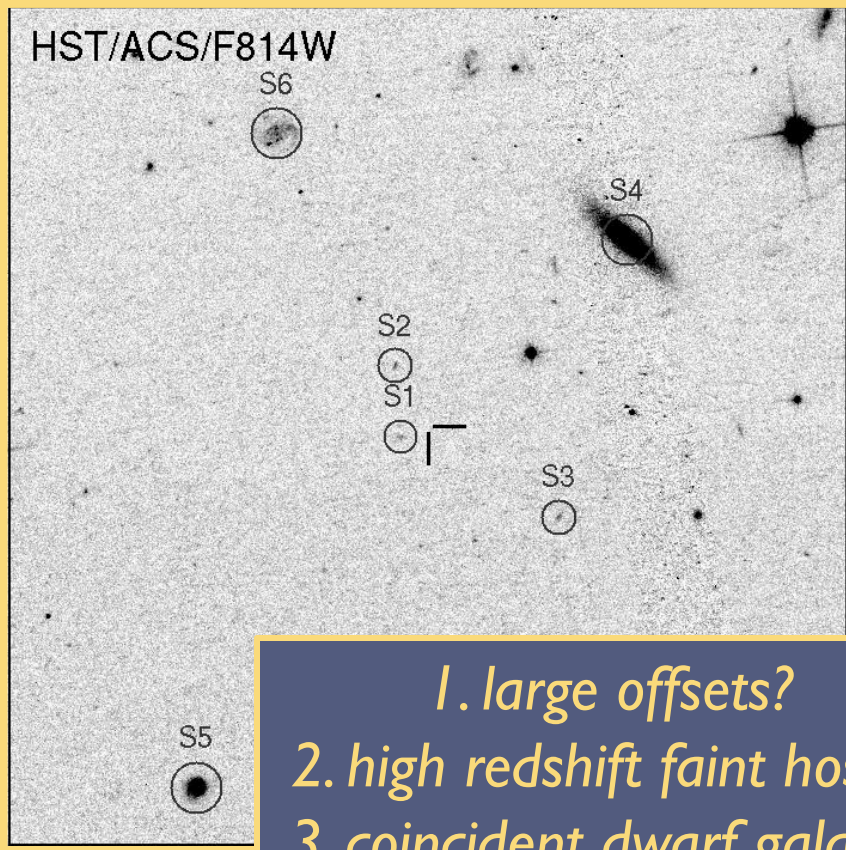


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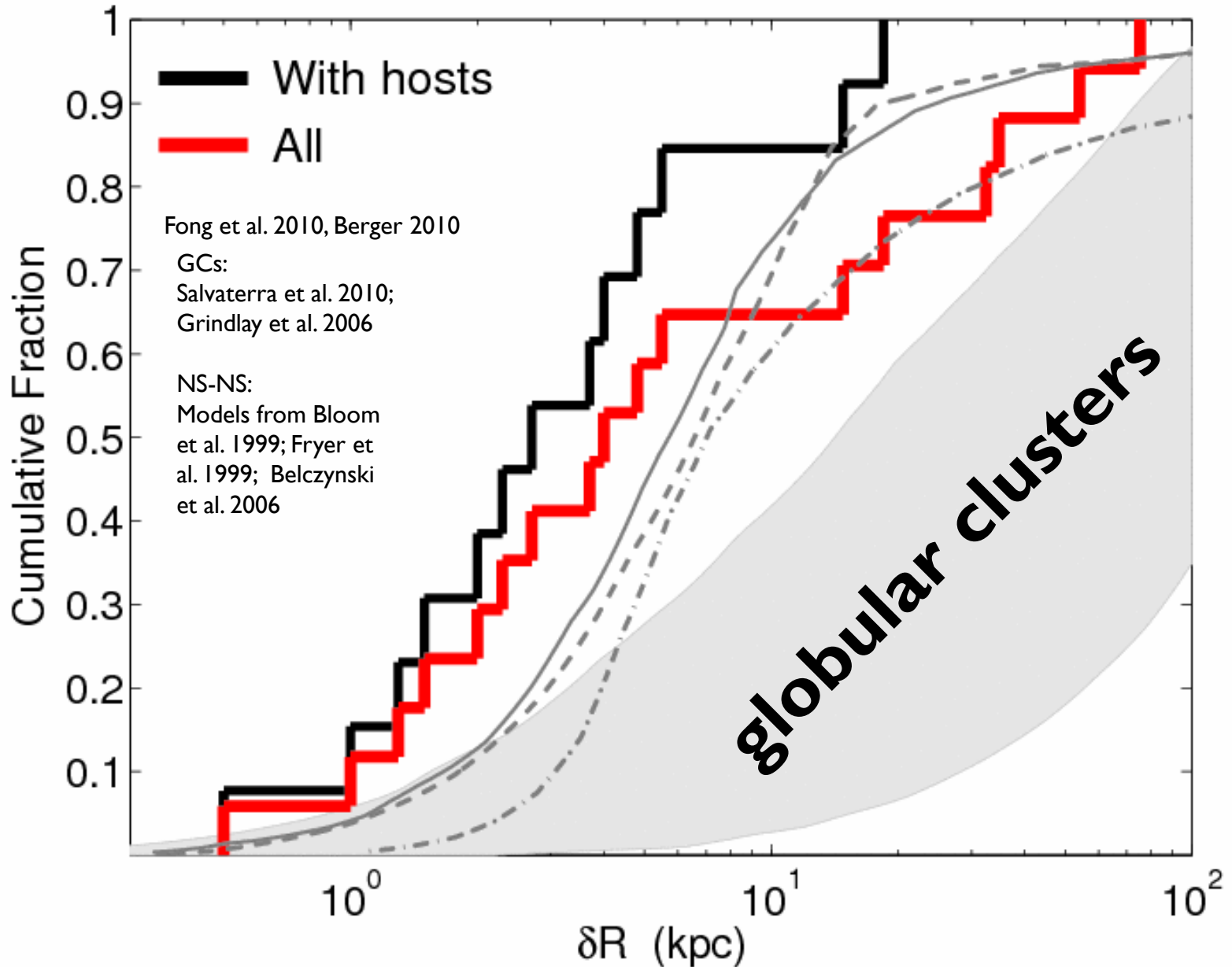
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1. large offsets?
2. high redshift faint hosts?
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Strong evidence for a highly-kicked progenitor system?

Short GRB offsets



The story so far...

● **Geometry of outflow?**

- Fraction are highly collimated
- Rates may be comparable with NS-NS predictions

● **Sub-parsec environment?**

- **LOW** densities, $\sim 10^{-2}-10^{-3}\text{cm}^{-3}$
- Median energy scale of 10^{51} erg

● **Nature of the progenitor?**

- Offsets provide best agreement with DNS models to date