Classification of variable stars in the WFCAM Transit Survey

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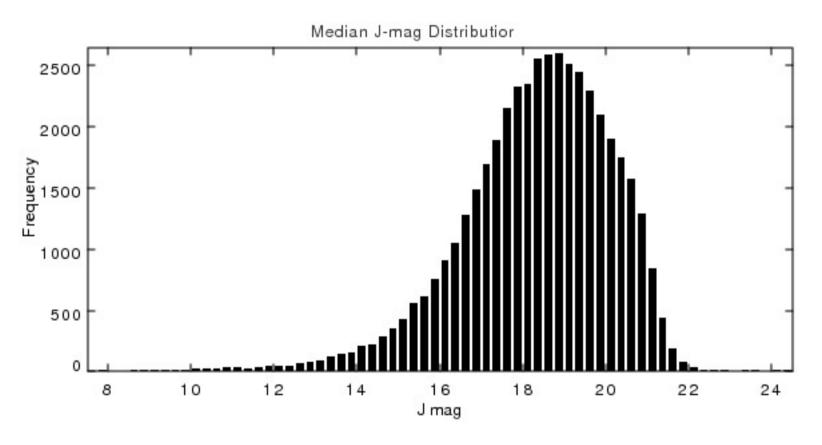
Outline

- Context
- Reduction of the WTS light curves
- Search for significant frequencies: bossirr and freq
- Classification of variable stars in the WTS 19hr field
- Summary and outlook

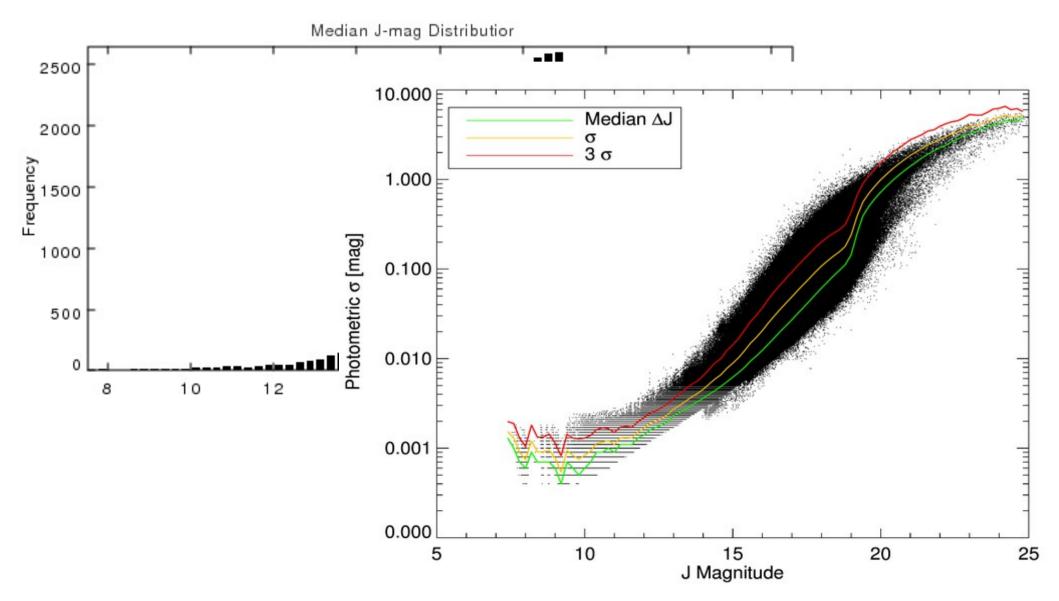
Variable stars are exciting!

- One of the principal areas of astronomical research
- In recent years time-resolved photometric data has marked a boom: OGLE, ASAS, Pan-STARRS, CoRoT, Kepler...
- Soon to come Gaia, VISTA
- All those new variables need to be classified!

Data Characterisation and Reduction of the WTS Light Curves



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Data Characterisation and Reduction of the WTS Light Curves

- Reject any measurement outside the $\langle J \rangle + 3\sigma(J)$ interval on a light curve by light curve basis
- Reject non-stellar objects, blended objects and objects fainter than the flux limit set in header
- Reject objects fainter than $\langle J \rangle = 19$ mag and with less than 20 data points

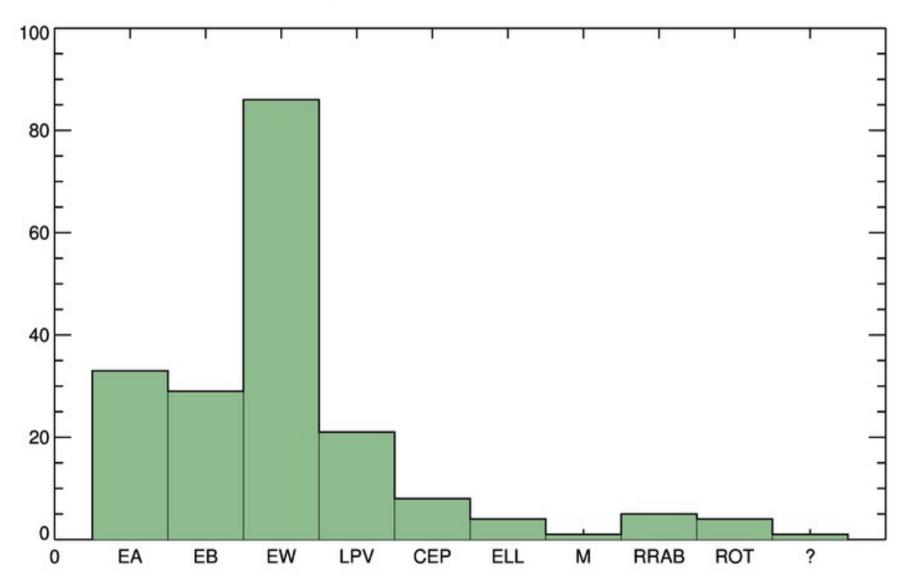
bossirr and freq

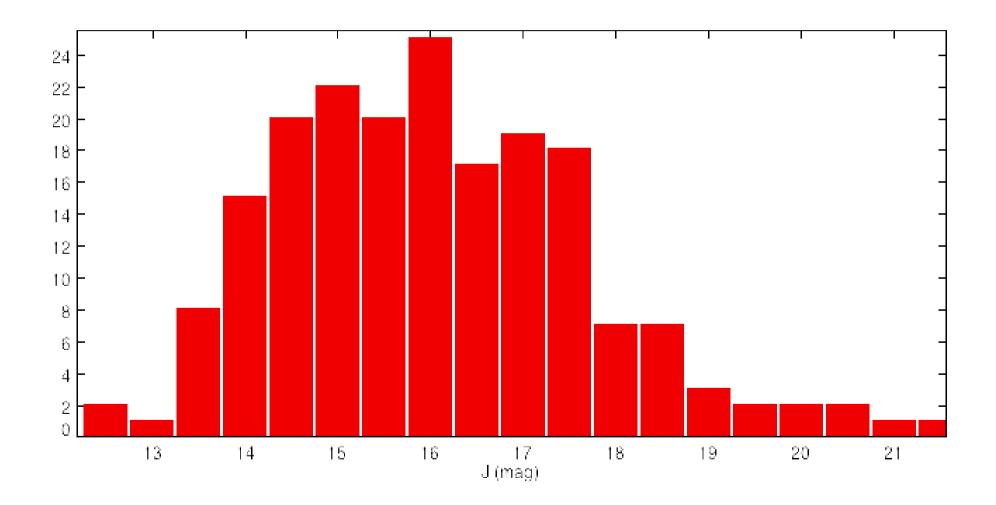
- bossirr performs non-linear sinusoidal fits to detect up to two significant frequencies
- *freq, by* least-square fitting, determines with greater precision the detected frequencies and obtains up to four harmonic modes of the form

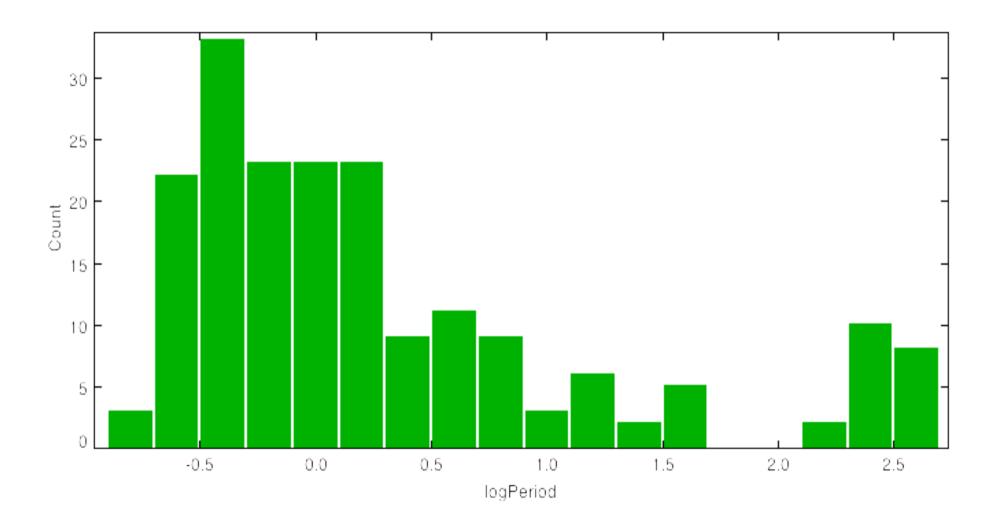
$$f(t) = \sum_{i=1}^{2} \sum_{j=1}^{4} A_{ij} \sin(2\pi f_i j t + \varphi_{ij}) + b_0$$

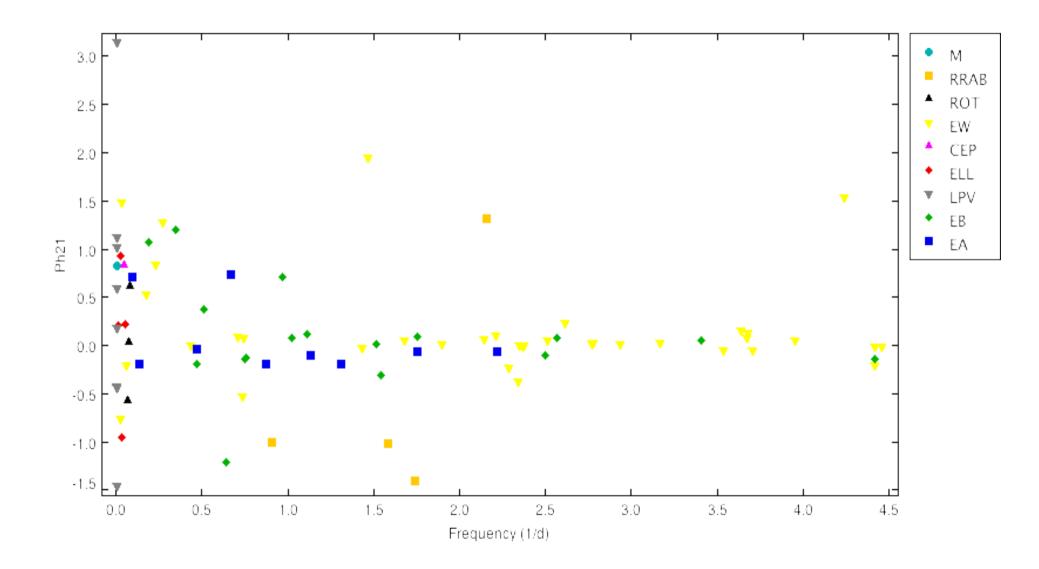
to determine their amplitudes and phases

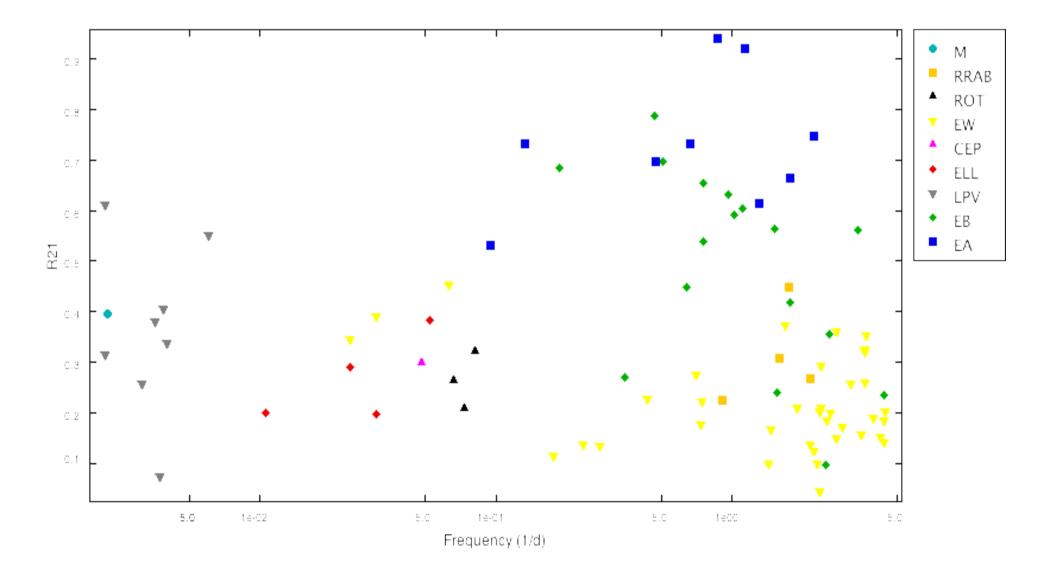
Classification

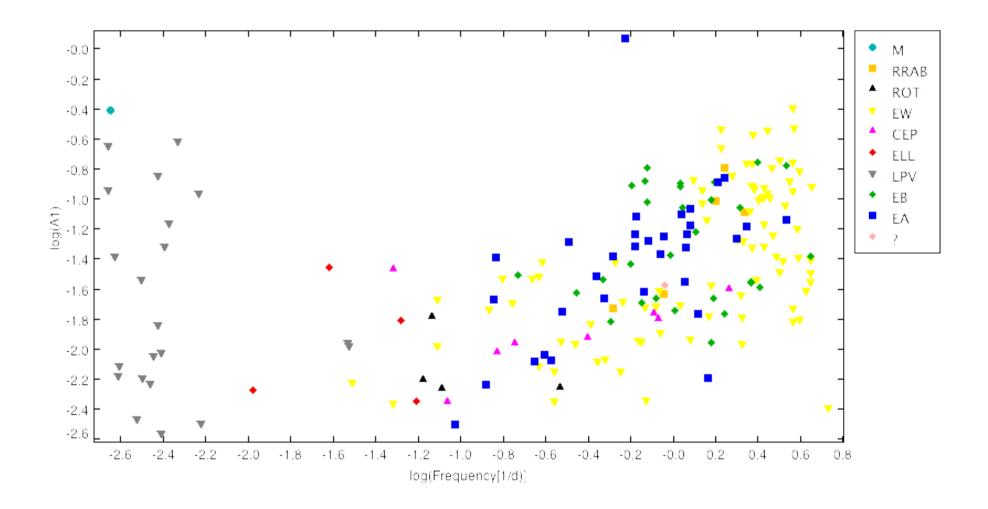


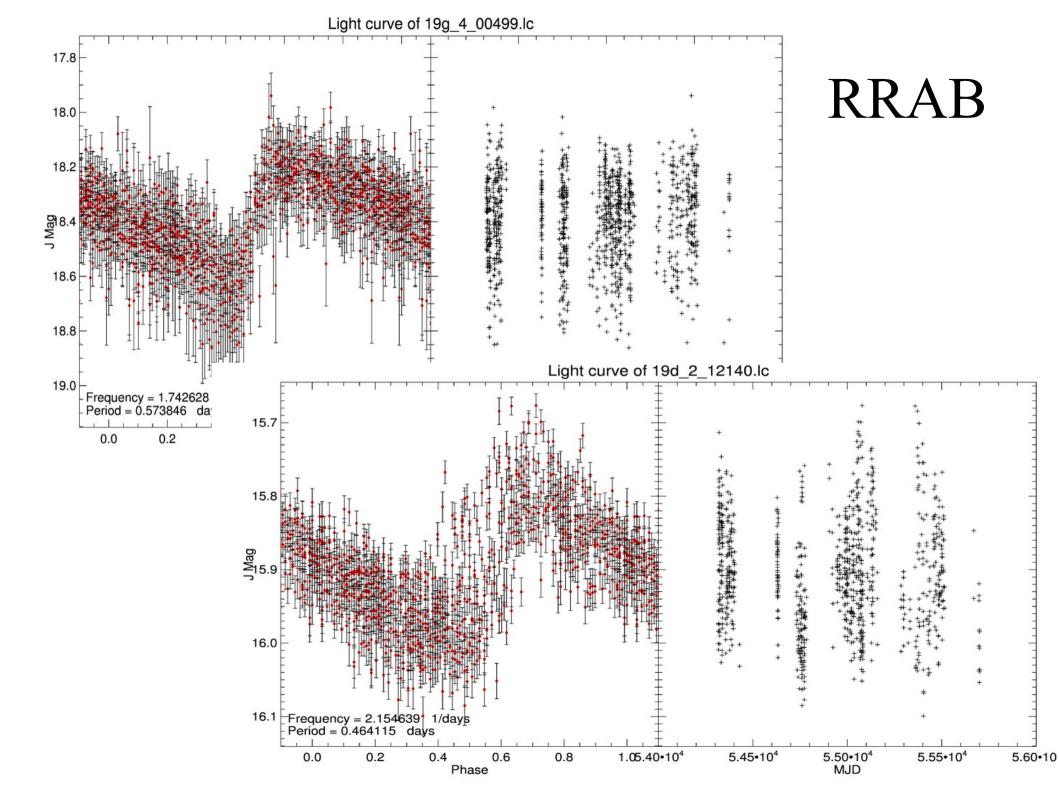


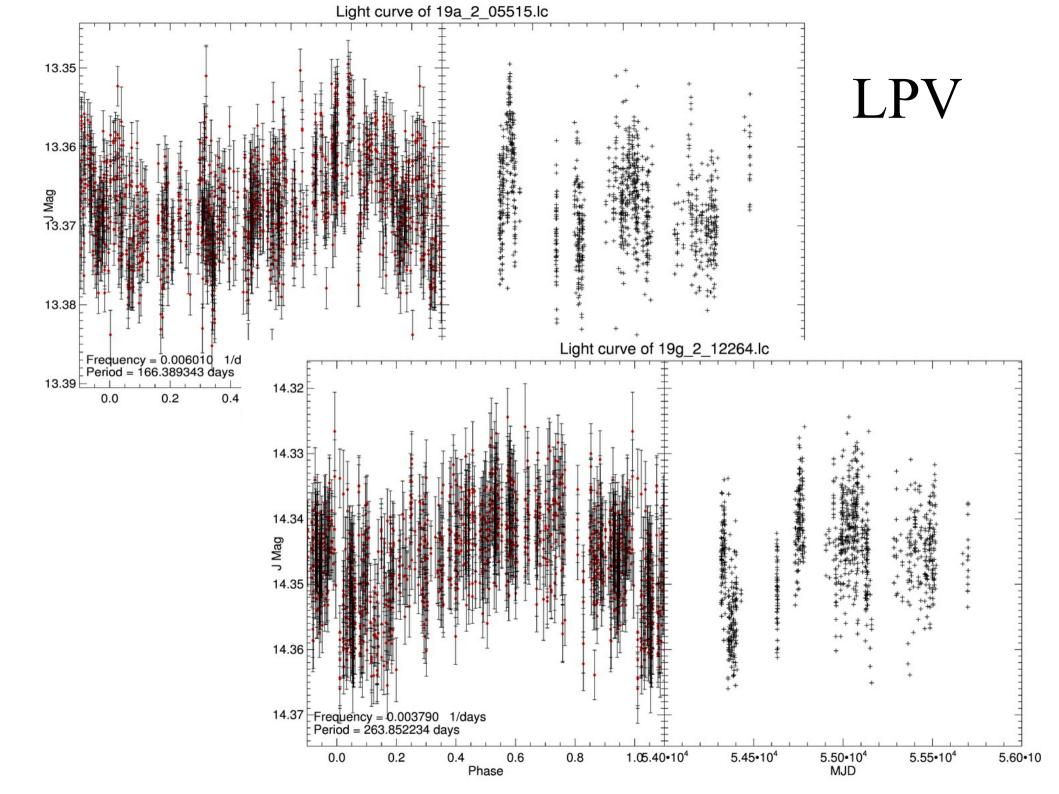


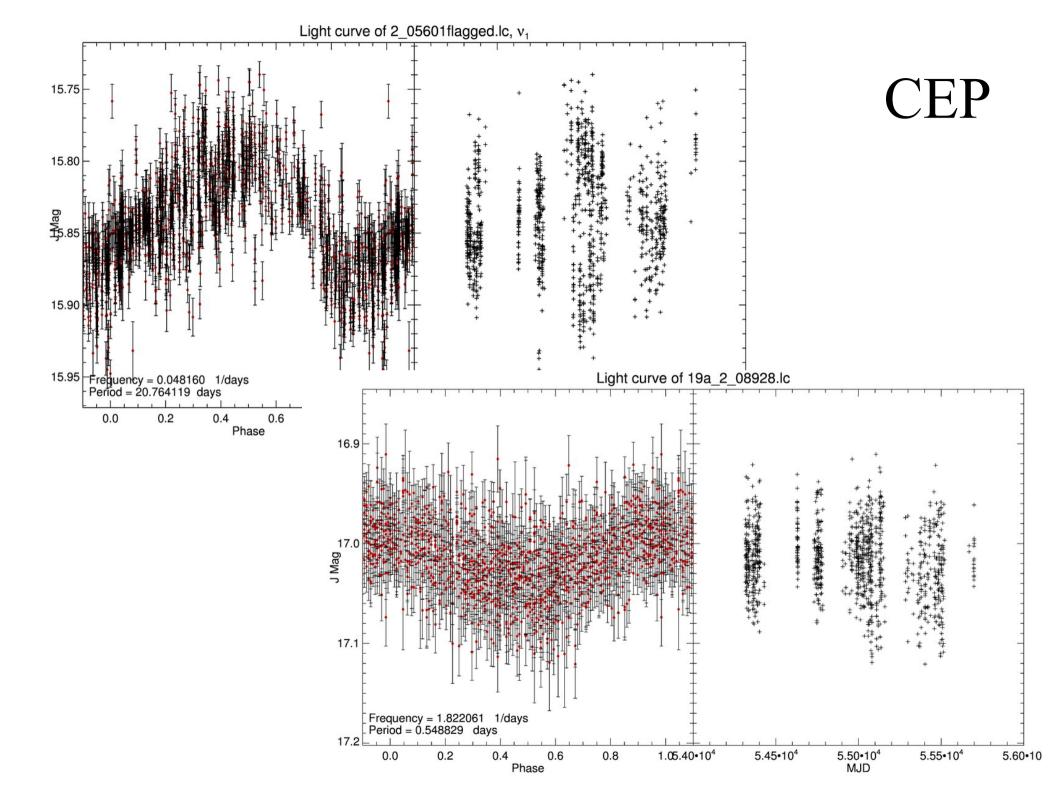


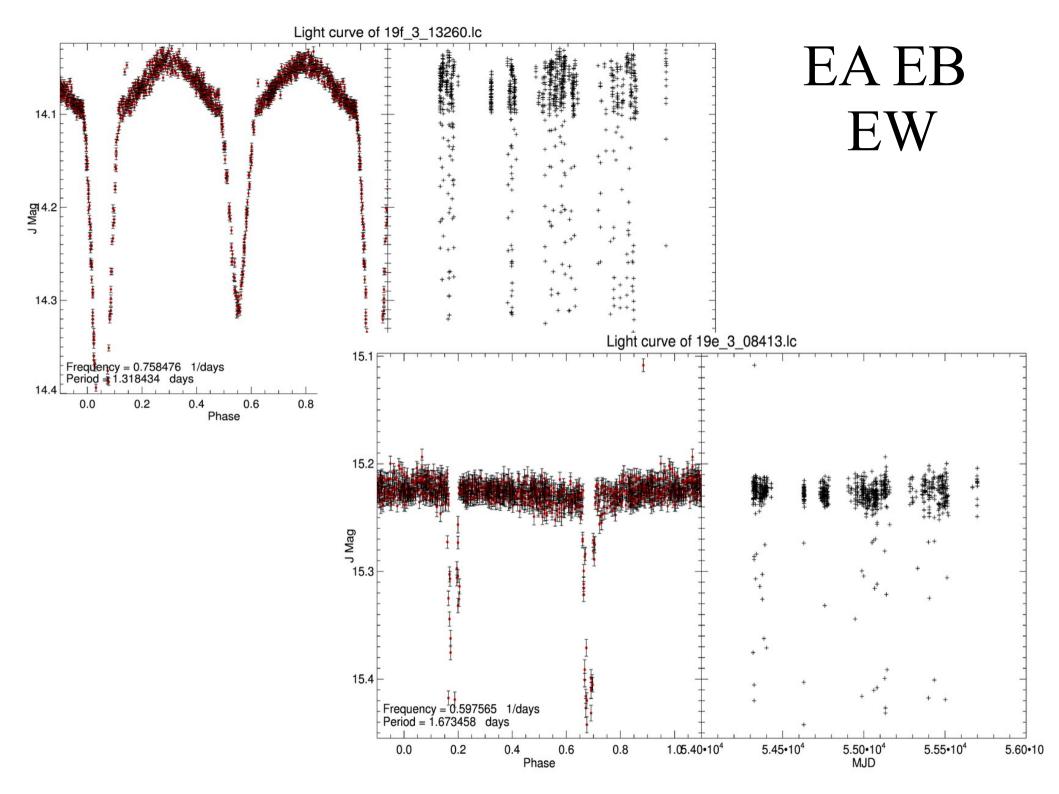


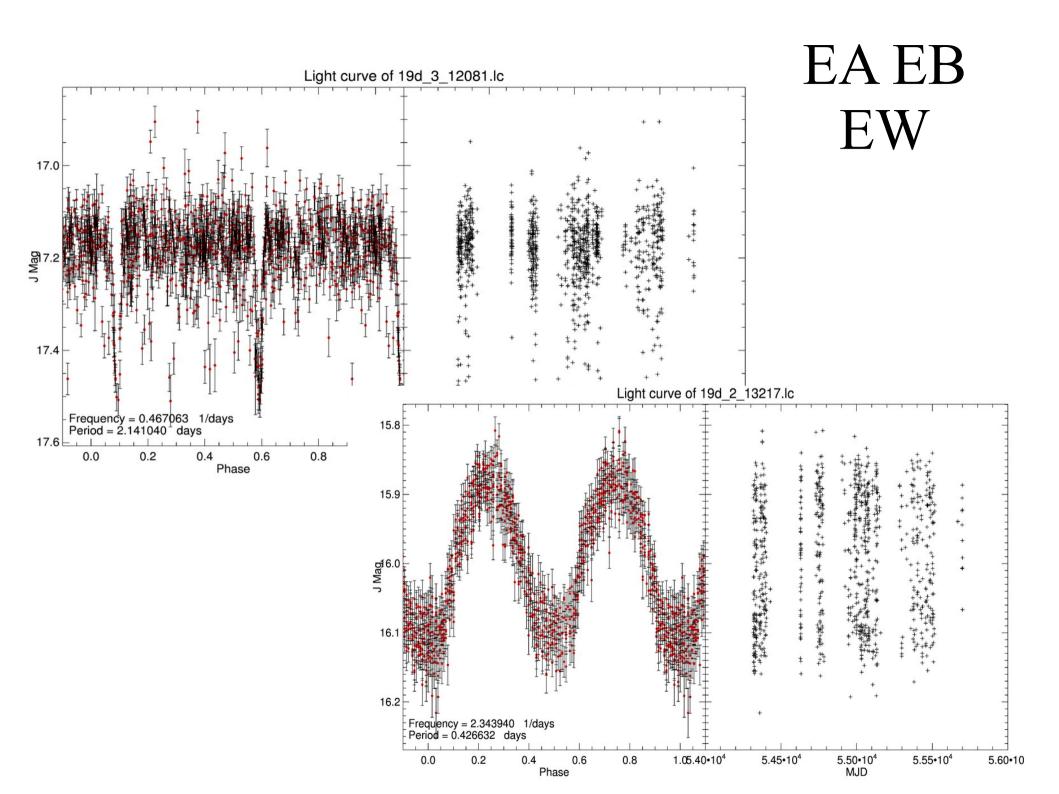












Summary

- 185 new classified variable stars detected
 - Another seven EWs already in VSX catalogue
- To be used for the classification of the rest of the WTS fields
- WTS to produce a complete training set for variable star classification in the nIR

Thank you!

