

Nucleocosmochronology and the r Process

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The basic program

- We know the solar system's initial abundance of $^{238}\text{U}/^{232}\text{Th}$:
 - $N(^{238}\text{U}/^{232}\text{Th}) = 0.431$
- We compute the production ratio $P(^{238}\text{U})/P(^{232}\text{Th})$ in the r-process
- We determine the age of the Galaxy by computing the age of the elements from these data

Single-Event Age

$$\frac{N(^{238}\text{U})}{N(^{232}\text{Th})} = \frac{P(^{238}\text{U}) e^{-\lambda_{238}T}}{P(^{232}\text{Th}) e^{-\lambda_{232}T}}$$

thus

$$\Delta^{\max} = \frac{1}{\lambda_{238} - \lambda_{232}} \ln \left(\frac{P(^{238}\text{U}) / P(^{232}\text{Th})}{N(^{238}\text{U}) / N(^{232}\text{Th})} \right)$$

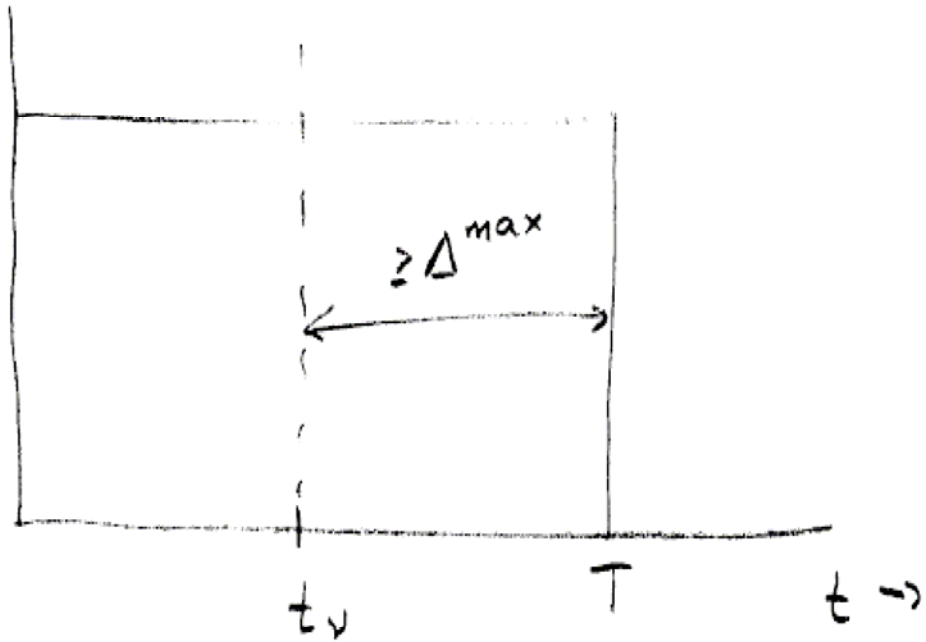
Chemical Evolution

$$T - t_{\nu} \geq \Delta^{\max}$$

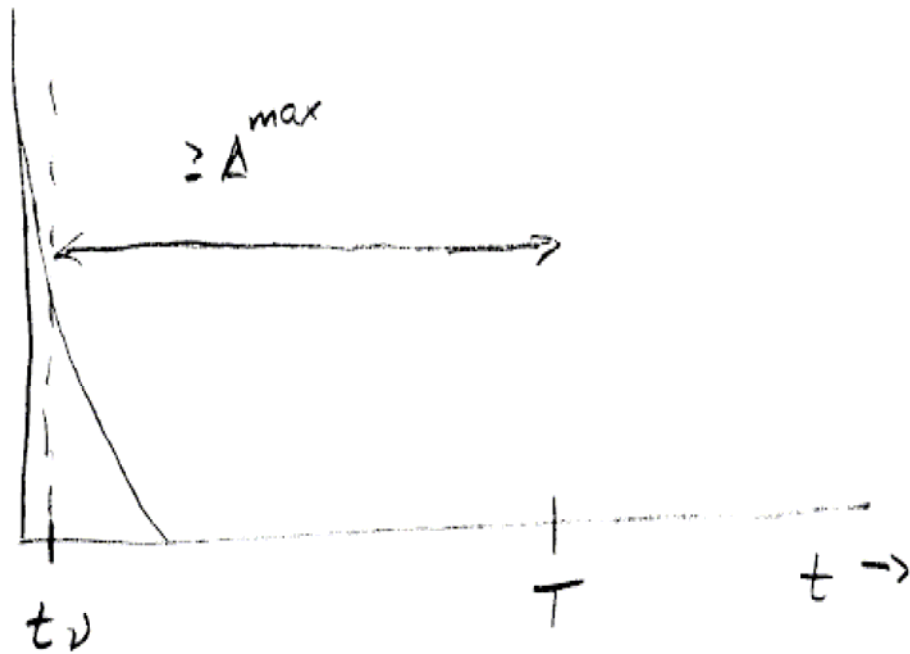
Schramm and Wasserburg (1970)

Meyer and Schramm (1986)

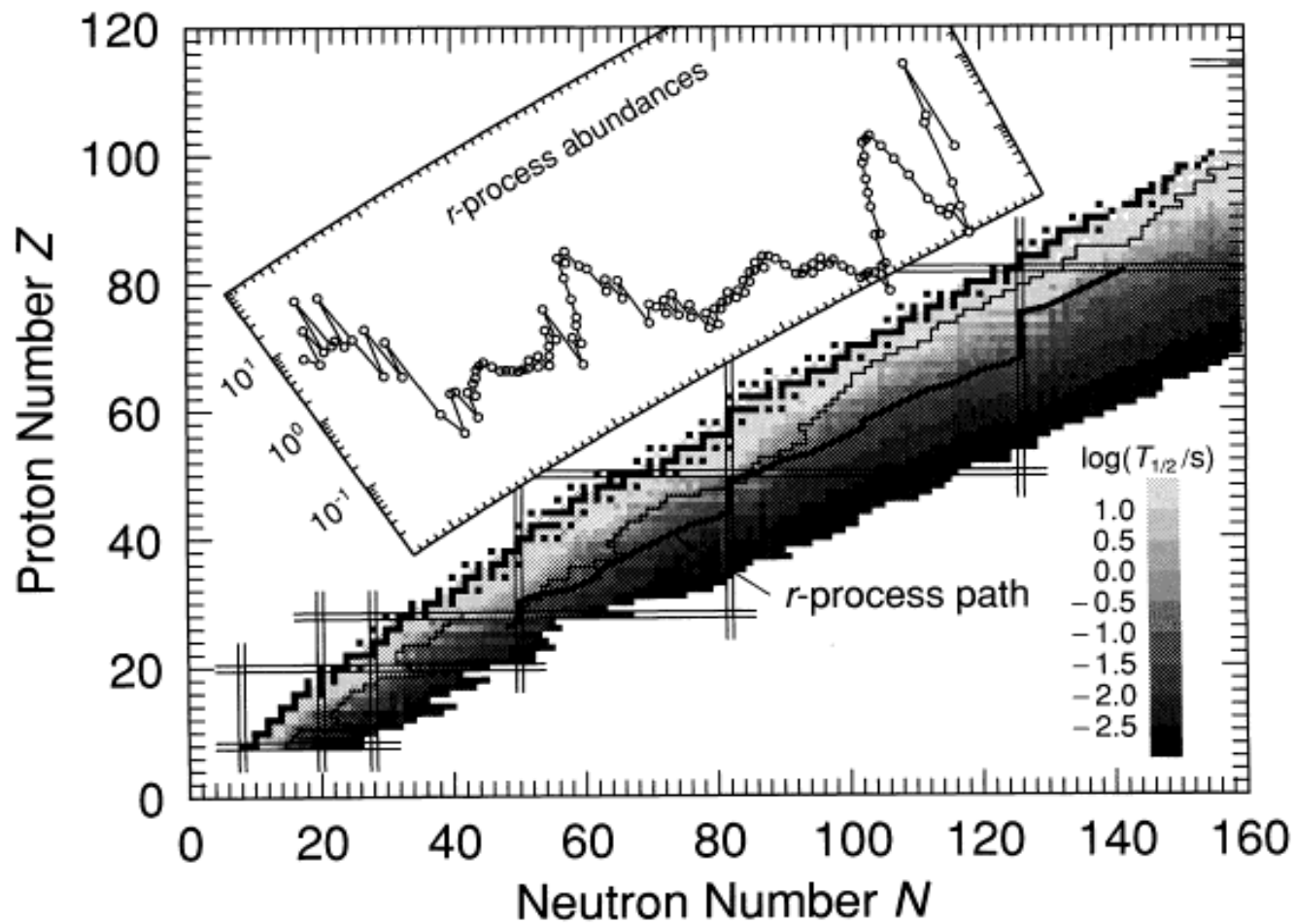
Effective
Nucleosyn.
Rate



Effective
Nucleosyn.
Rate



$$T \approx \Delta^{\max}$$



Production Ratio

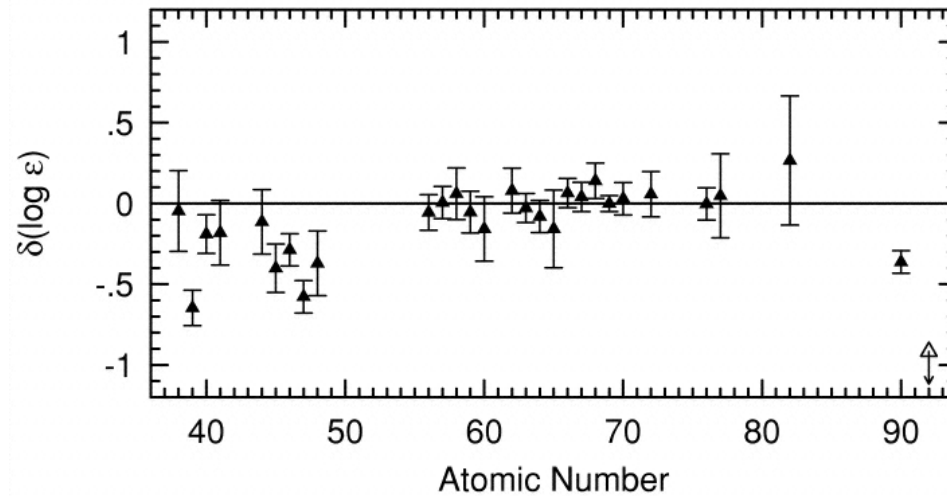
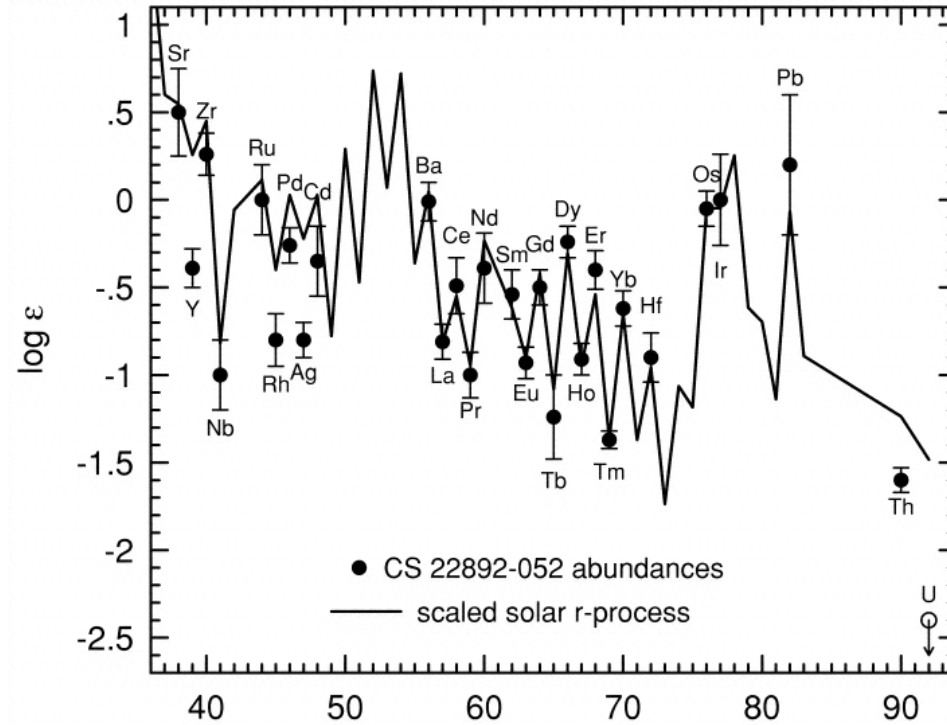
- Computed from models
- Simplest: equal production of progenitors
 - ^{238}U : ^{238}U (100%), ^{242}Pu (100%), ^{246}Cm (99.97%), and ^{250}Cm (18%) = 3.2
 - ^{232}Th : ^{232}Th (100%), ^{236}U (100%), ^{240}Pu (99.97%), ^{244}Pu (99.88%), ^{248}Cm (91.61%), ^{252}Cf (96.91%) = 5.8
progenitor
 - $P = 3.2/5.8 = 0.55 \Rightarrow T = 2.3 \text{ Gyr} \Rightarrow T_{\text{G}} > 6.9 \text{ Gyr}$
- R-Process models give P in the range 0.4 – 0.7.

Difficulties with Nucleocosmochronology

- Production ratio
 - Mechanism of the r-process itself
 - Uncertain nuclear physics
- Chemical evolution model
 - Factors of two uncertainties

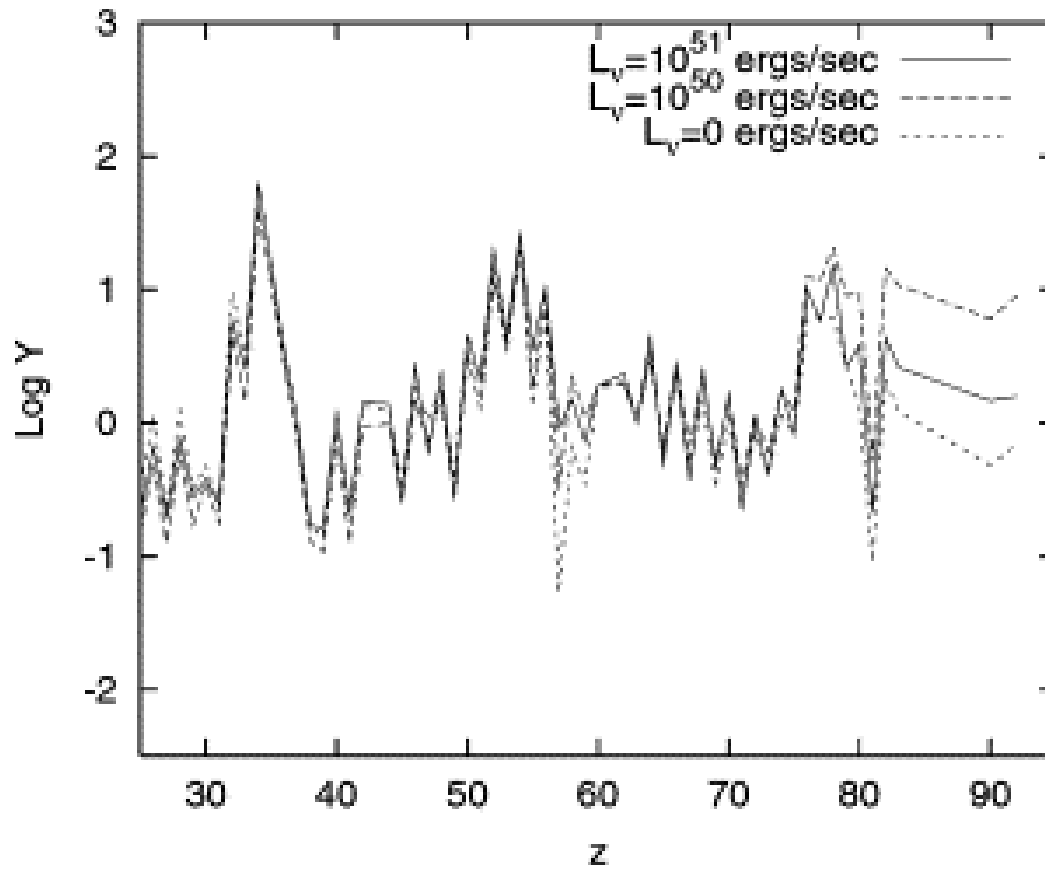
$$\Rightarrow T_{\text{Galaxy}} = 10 - 20 \text{ Gyr}$$

Sneden et al.
(2000)



Th/Eu=15.2 Gyr (+4.5-3.7)

Meyer and Truran (2000)

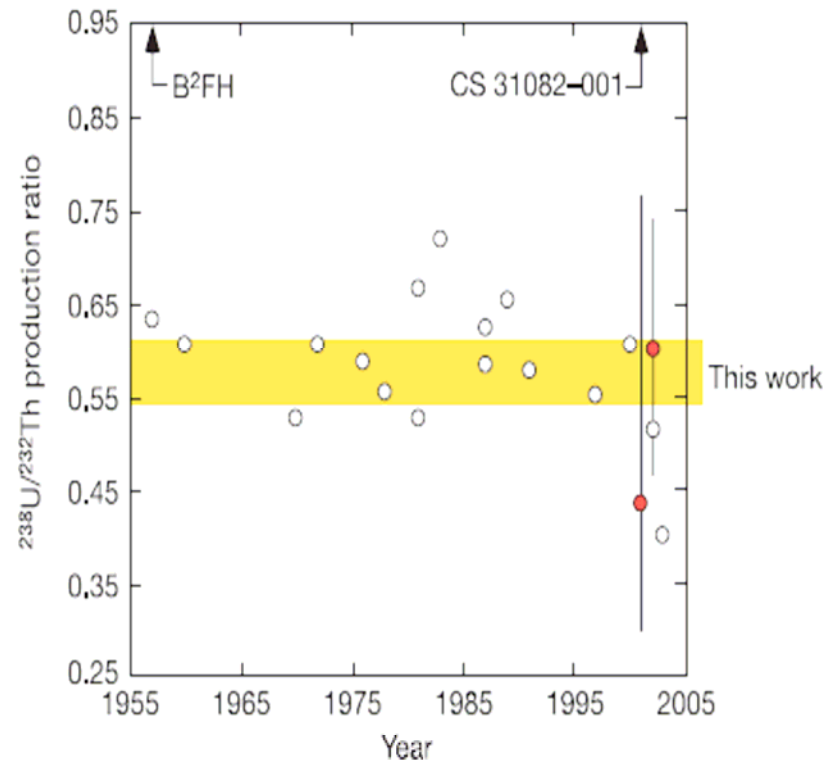


Otsuki, Mathews, and Kajino (2003)

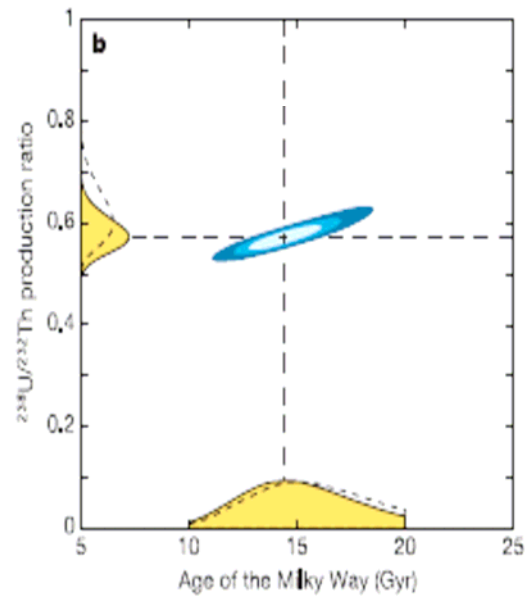
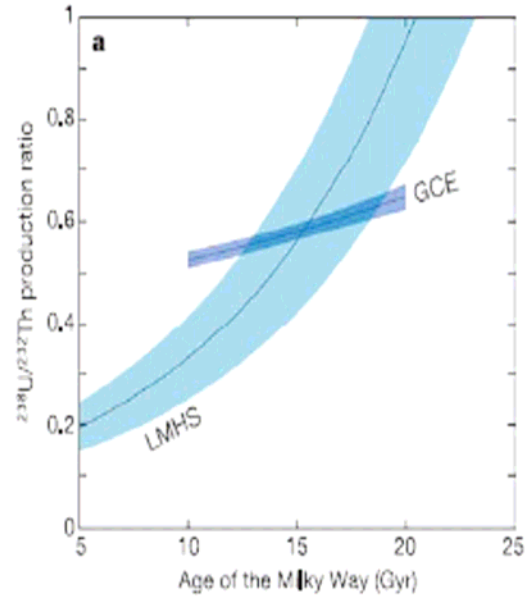
What about U/Th?

- Neighboring isotopes
- Some recent determinations:
 - Goriely and Arnould (2001) give $0.435^{+0.329}_{-0.137}$
 - Schatz et al. (2002) give $0.603^{+-0.139}$

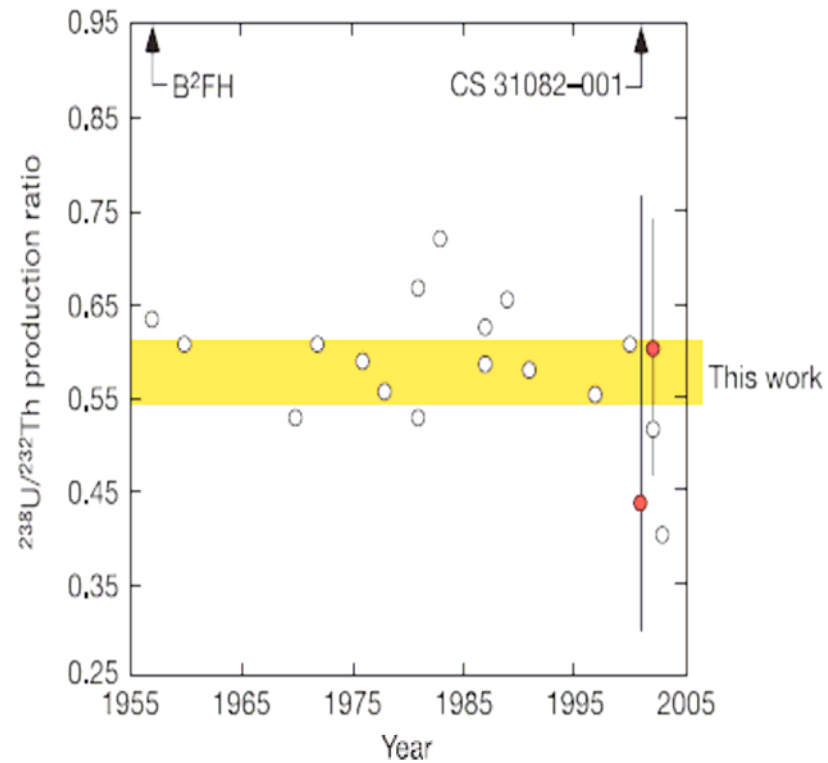
Dauphas (2005)



Dauphas (2005)



Dauphas (2005)



$$P = 0.571 + 0.037 - 0.031$$
$$\Rightarrow T_{\text{Gal}} = 14.5 + 2.8 - 2.2 \text{ Gyr}$$

Conclusions

- From metal-poor stars and chemical evolution, it is becoming possible to constrain the U/Th production ratio
- This will constrain r-process models and/or the nuclear physics input

Radioactivity with Astronomies