Optimal Binning Kaastra & Bleeker, 2016, A&A 587, A151

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Introduction

We had

$$n_{ph}(c) = \int_0^\infty R(c, E) \cdot A(E) \cdot F(E) \, dE + n_{background}(c) \tag{1}$$

and discretized to

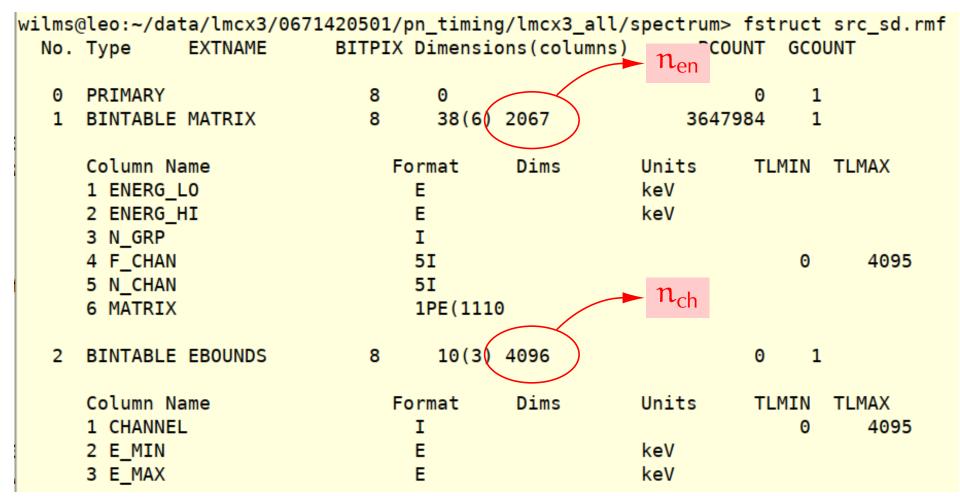
$$\begin{split} N_{ph}(c) &= \Delta T \cdot \sum_{i=0}^{n_{ch}} A(E_i) \cdot R(c,i) \cdot F(E_i) \cdot \Delta E_i + N_{background}(c) \\ &\quad \forall c \in \{1,2,\ldots,n_{en}\} \ (2) \end{split}$$

but

- What value should we choose for the number of energy channels, n_{en} ?
- What value should we choose for the number of spectral (PHA/PI) channels, n_{ch} ?
- \implies Optimal Binning

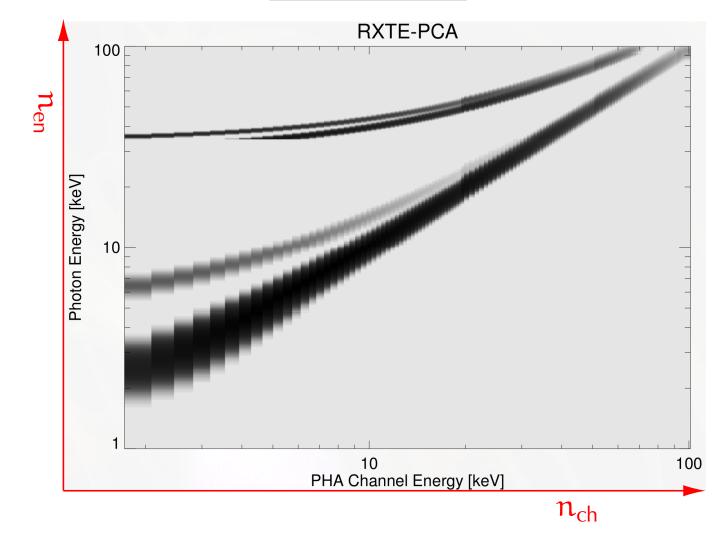
Introduction

Implementation in FITS:



Note: RMF does not have to be square, i.e., $n_{ch} \neq n_{en}$

Introduction



Optimum choice of n_{ch} and n_{en} : Kaastra & Bleeker, 2016, A&A 587, A151 this presentation follows the paper very closely

In simplified notation, ignoring ARF for the moment: measured flux

$$s(E') = \int_0^\infty R(E', E)f(E)dE$$
(3)

Analytical integration impossible \implies discretize ("model energy grid" or "model grid", boundaries E_{1j} , E_{2j} , center $E_j = (E_{1j} + E_{2j})/2$, width $\Delta E_j = E_{2j} - E_{1j}$ Then:

$$S_{i} = \sum_{j} R_{ij} \cdot F_{j} \text{ where } F_{j} = \int_{E_{1j}}^{E_{2j}} f(E) dE$$
(4)

To evaluate, however, most analysis packages use

$$F_{j} = f(E_{j})\Delta E_{j}$$
(5)

 \implies implies $f(E) \sim$ const. over ΔE_j .

⇒ use "sufficiently large number of bins"

(e.g., Suzaku RMF: 1 eV, XMM-EPIC-pn: 5 eV)

Implies O(10000) bins for typical Si-based detectors today

What is "sufficiently large"?

Determination of narrow line centroid:

accuracy of E_{line} : σE

$$\propto \frac{\sigma}{\sqrt{N}}$$

(6)

where σ : resolution, N: photon number

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Determination of narrow line centroid:

accuracy of E_{line} : σ

$$E\propto rac{0}{\sqrt{N}}$$

(6)

where σ : resolution, N: photon number

• Athena-WFI ($\sigma \sim 20 \text{ eV} \ 1 \text{ keV}$), line w/20000 counts: $\sigma E = 0.1 \text{ eV}$ $\implies 100\,000 \text{ bins}$

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- Athena-WFI ($\sigma \sim 20 \text{ eV} \ 1 \text{ keV}$), line w/20000 counts: $\sigma E = 0.1 \text{ eV}$ $\implies 100\,000 \text{ bins}$
- Athena-X-IFU ($\sigma \sim 1 \text{ eV}$ @ 1 keV), line w/20000 counts: $\sigma E = 0.005 \text{ eV}$ $\implies 2000000 \text{ bins}$

Just increasing number of n_{ch} will not help with future detectors.

Already now model evaluations are extremely expensive, e.g., NuSTAR and Suzaku require rebinning of matrix.

K&B: Optimize evaluation of F_j

Assuming Gaussian redistribution

• Oth order (... the XSPEC way)

$$f_{1,0}(E) = N\delta(E - E_j)$$
 (7)

where N: number of photons in bin

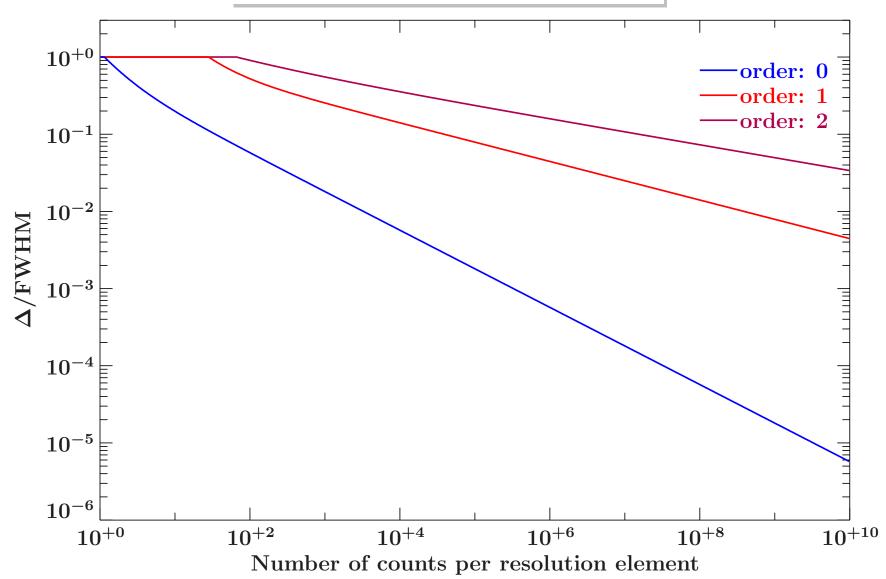
• 1th order: use average photon energy:

$$f_{1,1}(E) = N\delta(E - E_a)$$
 with $E_a = \int_{E_{1j}}^{E_{2j}} f(E)EdE$ (8)

• 2nd order: consider average and variance of photon energy (1st and 2nd moments):

$$f_{1,2}(E) = N \exp\left((E - E_a)/2\tau^2\right)$$
 with $\tau^2 = \int_{E_{1j}}^{E_{2j}} f(E)(E - E_a)^2 dE$ (9)

Number of energy channels



after K&B, Fig. 1

Already using 1st order gives order of magnitude improvement in number of channels!

K&B:

- recommend 1st order: good compromise between numerical complexity and improvement
- They also present equations to take into account ARF effects

implemented in SPEX and ISIS, but in principle also doable for XSPEC with a binfile for ftrbnrmf

Number of PI channels

Number of PI channels: How should I bin my data? Information theory: Shannon information theorem: for a band limited function

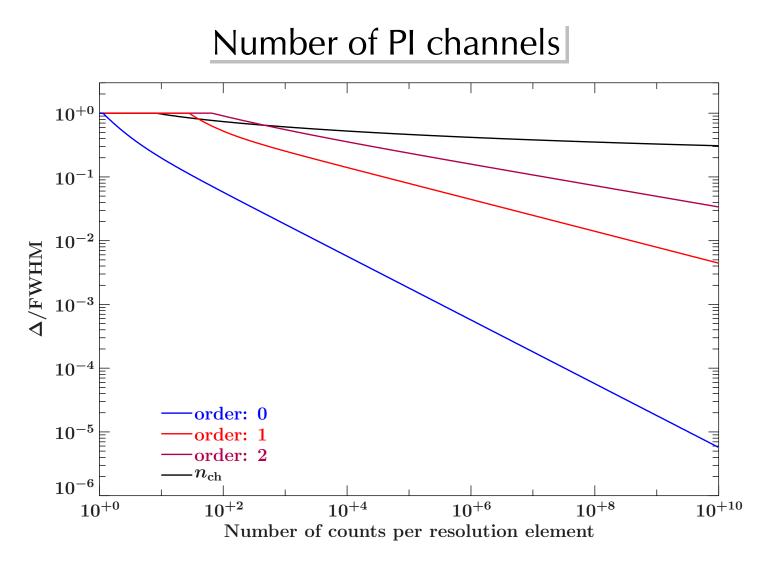
$$f(x) = \sum_{n = -\infty}^{+\infty} f(n\Delta) \frac{\sin \pi (x/\Delta - n)}{\pi (x/\Delta - n)}$$
(10)

where $\Delta = 1/2W$ and where W: frequency above which Fourier transform of f(x) is 0.

i.e., can describe f(x) fully using its values at discretely sampled values

X-rays: measure channel integrated quantities at boundaries $m\Delta \implies$ Integrate the above:

$$F(m\Delta) = \frac{\Delta}{\pi} \sum_{n=-\infty}^{\infty} f(n\Delta) \left(\frac{\pi}{2} + Si[\pi(m-n)]\right)$$
(11)



after K&B, Fig. 1

Determine Δ optimal for N counts per resolution element and R resolution elements:

$$\frac{\Delta}{\text{FWHM}} = \begin{cases} 1 & \text{if } x \leq 2.119\\ \frac{0.08 + 7/x + 1.8/x^2}{1 + 5.9/x} & \text{otherwise} \end{cases}$$
(12)

where $x = \ln (N(1 + 0.2 \ln R))$.