# A Exercises 1

## A.1 Magnitude of 4-velocity

Calculate the magnitude (vector length) of the 4-velocity u, defined by

$$u^{\mu} = \frac{dx^{\mu}}{d\tau}$$

- Use the definition of the 4-coordinates.
- Since this is a scaler, you can choose to work in any inertial frame choose the easiest one.

# A.2 4-acceleration

Show that the 4-acceleration is perpendicular to the 4-velocity w.r.t to the metric defined

$$a^{\mu} = \frac{d}{d\tau} \frac{p^{\mu}}{m_0}$$

## A.3 Coordinate transformation to spherical coordinates

Calculate the metric tensor of Minkowski space-time in spherical coordinates, defined by

$$x = r \sin \theta \sin \phi$$
$$y = r \sin \theta \cos \phi$$
$$z = r \cos \theta$$

- Calculate the differentials dx, dy, dz by deriving the transformations with respect to the new variables  $(r, \theta, \phi)$ . Note that  $x = x(r, \theta, \phi)$  - you will use the chain rule.
- Plug your results into the line element  $ds^2$  of the Minkowski space-time in Cartesian coordinates.
- By inspection of the result, find the metric tensor for Minkowski space-time in spherical coordinates.

## A.4 Coordinate transformation to elliptical coordinates

Calculate the metric of Minkowski space-time for the transformation to elliptical coordinates, defined by

$$x = \sqrt{r^2 + a^2} \sin \theta \sin \phi$$
  

$$y = \sqrt{r^2 + a^2} \sin \theta \cos \phi$$
  

$$z = r \cos \theta$$

For a = 0, this is the same as in exercise A.3. The path of calculation is the same.

#### A.5 Lorentz transformation

A Lorentz transformation is defined by  $(c\,t,x,y,z) \longrightarrow (c\,t',x',y',z')$  :

$$t' = \gamma \left( t - \frac{v_x}{c^2} x \right)$$
$$x' = \gamma (x - v_x t)$$
$$y' = y$$
$$z' = z$$

Show that a space-time interval |s'|, i.e. the vector length of 4-vector s', is the same as for the original vector 's' - i.e. that vector lengths are invariant under Lorentz transformations.

#### A.6 Keplerian angular velocity

Calculate for a circular orbit around mass M the angular velocity  $\omega$ , which is a function of radius.

# A.7 Effective potential & Keplerian orbits

Re-educate yourself on Keplerian orbits in a Newtonian potential, central mass M.

- Start from  $E = E_{kin} + E_{pot}$
- Express the velocity v in a radial and tangential component
- Introduce the angular momentum  $l = r^2 \dot{\phi}$
- Read off the form of the effective potential  $V_{\text{eff}}(r)$
- What are the turning points of the motion, as a function of E and l? (dr/dt = 0)
- In the energy equation, transform from  $\dot{r}$  to  $\frac{dr}{d\phi}$
- Transform once again to u = 1/r
- Take the derivative  $\frac{d}{d\phi}$
- The second order differential equation for  $u(\phi)$  obtained is "easy". Can you guess the solution?
- Verify that  $u(\phi) = 1 + e \cos \phi$  solves the equation
- What are the turning points, as a function of l and e?